

# WDMAM project and Compilation of global marine magnetic anomaly data

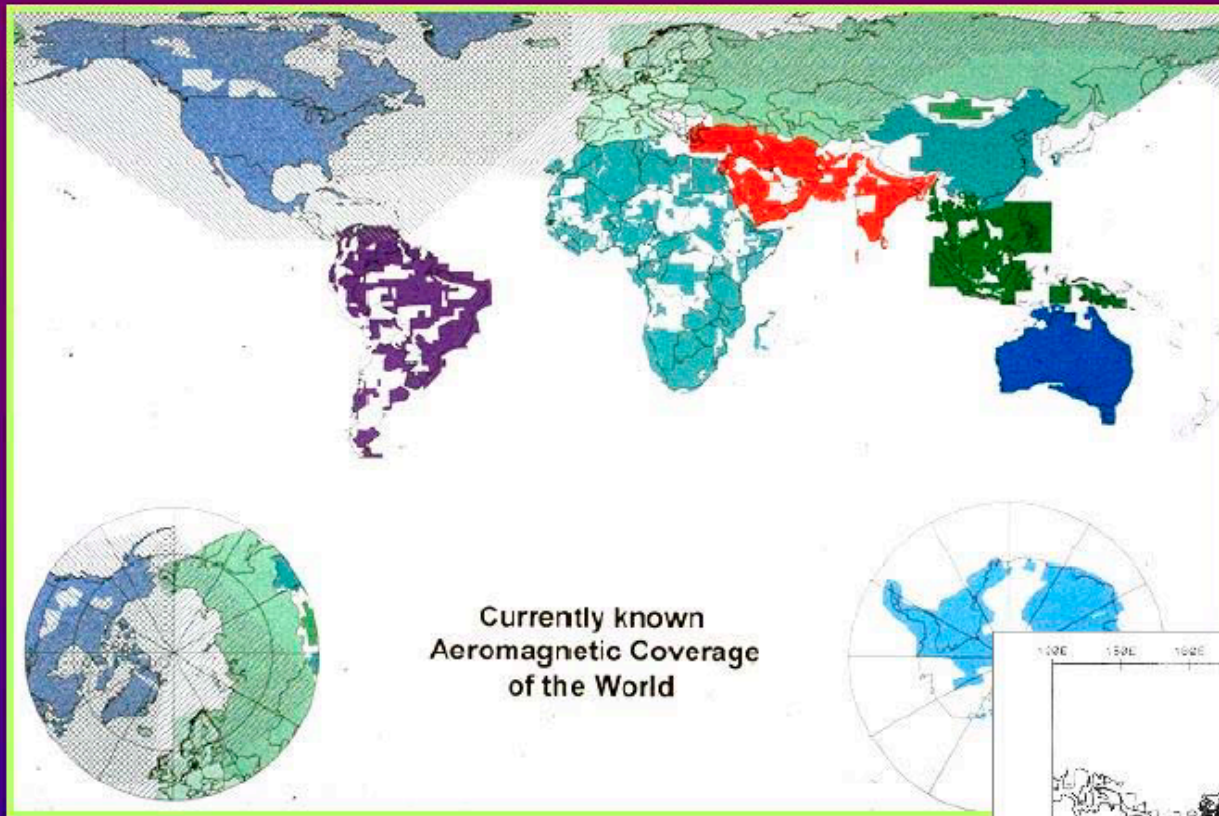
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# World Digital Magnetic Anomaly Map (WDMAM) project

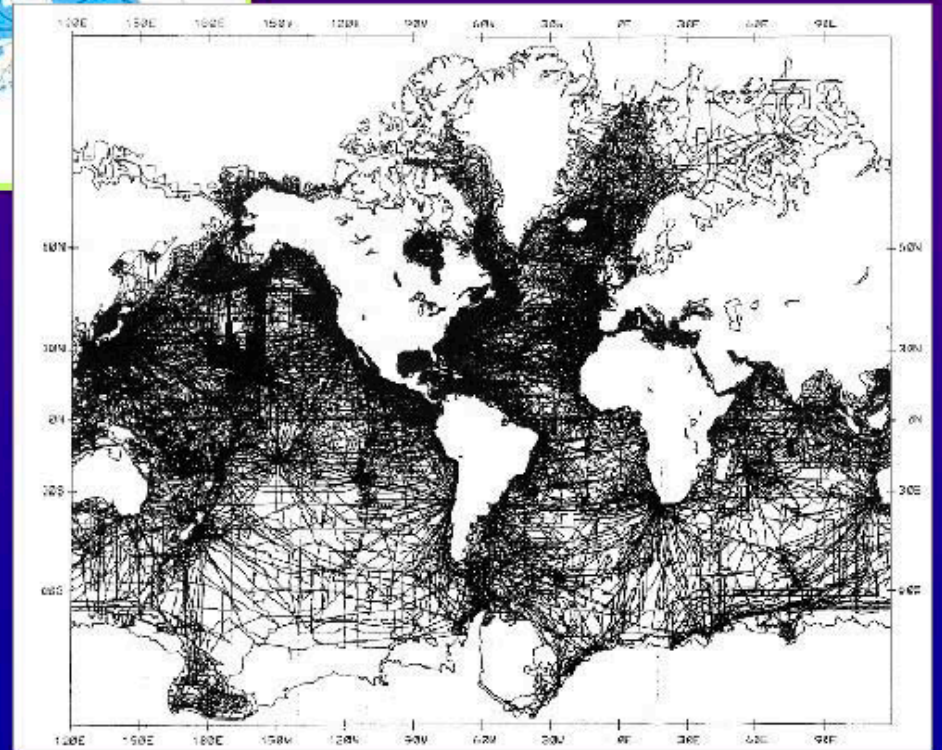


Currently known  
Aeromagnetic Coverage  
of the World

The patchwork of detailed  
aeromagnetic surveys  
over the continents...

...and the tracks of  
oceanographic survey  
vessels across the  
oceans...

...together make up an enormous  
but disparate body of data about  
global magnetic anomalies and, in  
turn, global geology.



- Reeves, Macnab & Maschenkov, 1998  
EOS, Vol 79, No. 28, p 338.

# WDMAM

- The WDMAM (World Digital Magnetic Anomaly Map) is an international scientific project under the auspices of IAGA (International Association of Geomagnetism and Aeronomy) and CGMW (Commission for the Geological Map of the World), aiming to compile and make available magnetic anomalies caused by the Earth lithosphere, on continental and oceanic areas, in a comprehensive way, all over the World.
- Various kinds of aeromagnetic, marine and satellite data are merged into a global magnetic anomaly data set.
- The project started in 2003.
- The 1<sup>st</sup> version of WDMAM was published in 2007.

**WDMAM, IAGA-Task Group Executives 2005-2007 (WDMAM 2007)**

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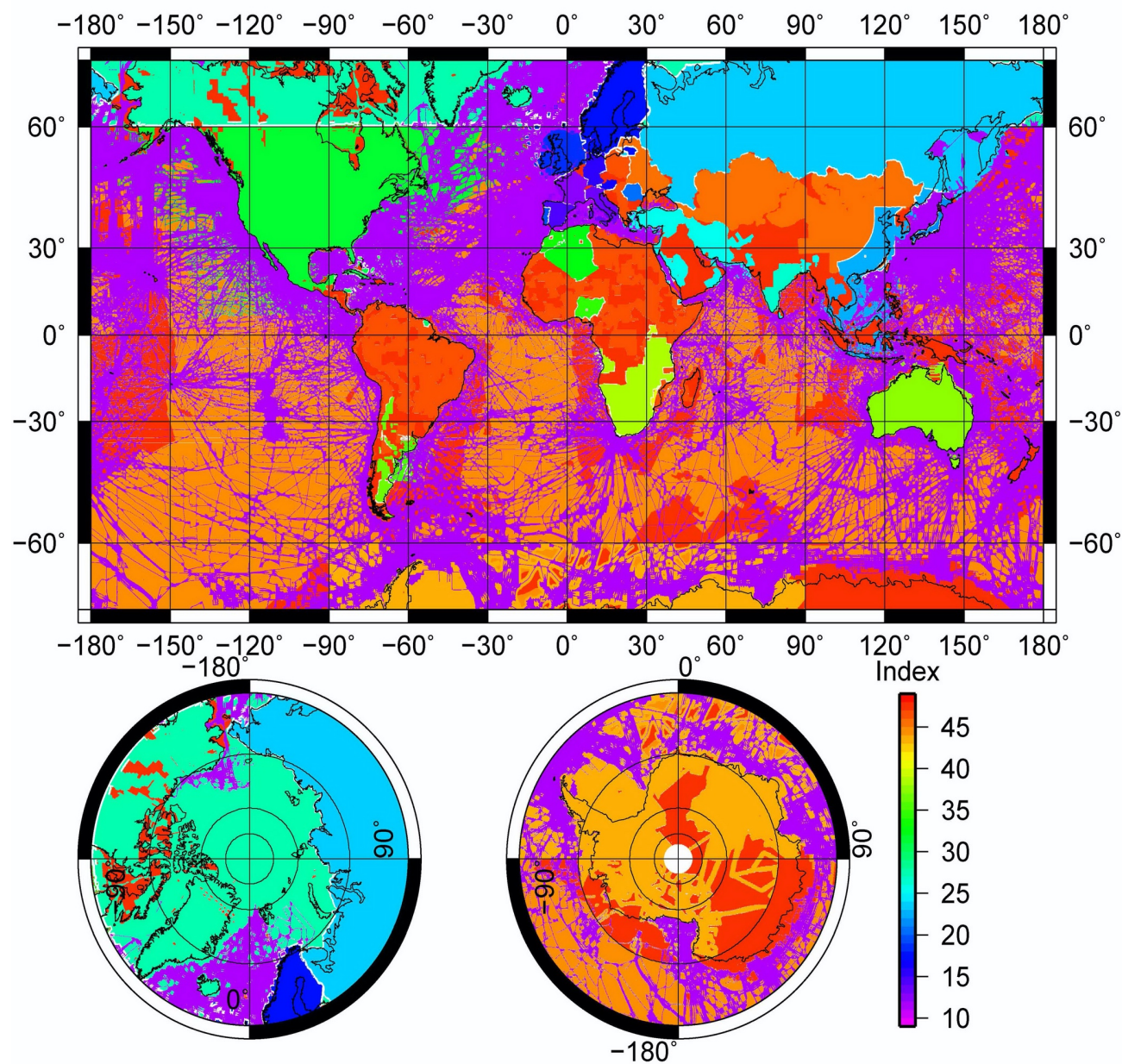




# WDMAM Version 2

- The 2<sup>nd</sup> version was published in 2015.  
WDMAM Task Force: J. Dyment (chair), M. Catalan (co-chair),  
A. de Santis, M. Hamoudi, T. Ishihara, J. Korhonen, V. Lesur,  
T. Litvinova, J. Luis, B. Meyer, P. Milligan, M. Nakanishi, S. Okuma,  
M. Pilkington, M. Purucker, D. Ravat, E. Thebault
- The compiled map (jpeg format) and grid data set are freely available at <http://www.wdmam.org>.
- Grid data set 3' x 3' (0.05° x 0.05°, about 5 km interval) grid
- Oceanic areas at altitude of 0 km  
marine track line data + magnetic anomaly model
  - Continental areas at altitude of 5 km  
merging of various kind of data sets
  - Long wavelength components (n=16 to 100) from satellite data
  - Satellite data where no aeromagnetic and marine data are available

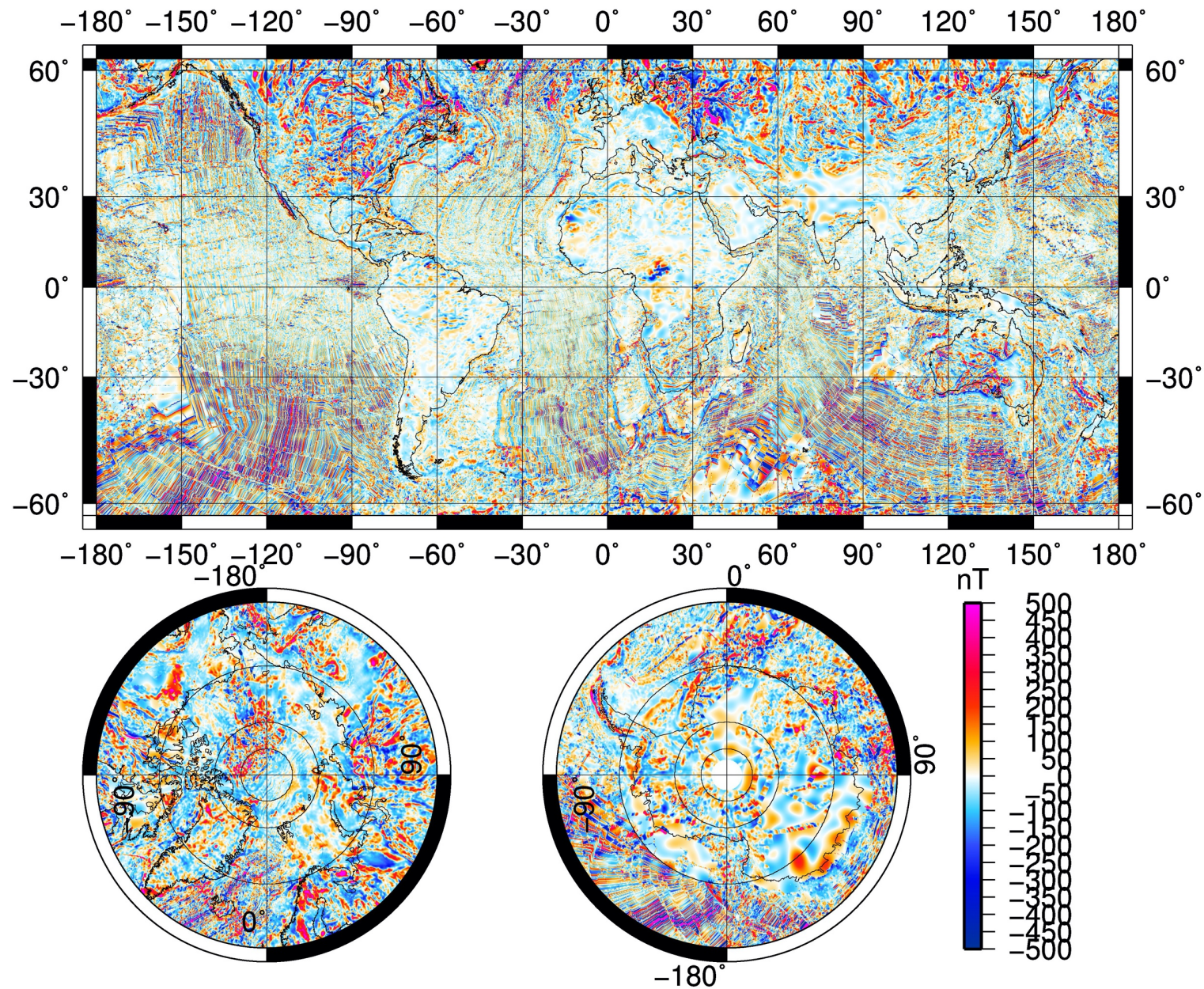
# Data sets 11 to 47 used in WDMAM version 2.0



11	Marine data set
12	France
13	Italy
14	Spain
15	Germany
16	Austria
17	Fennoscandia compilation
18	UK
19	Portugal
20	Romania
21	Japan
22	East Asia
23	Russia
24	Djibouti
25	Middle East
26	India
27	Circum Arctic Compilation
28	Guadeloupe
29	French Guiana
30	Bahama/Cuba
31	Nure/Namag
32	Algeria
33	Morroco
34	Nigeria
35	Argentina Shelf
36	Argentina Inland
37	Australia
38	South African Compilation
39	Uganda
40	Mozambique
41	SaNaBoZi (Compilation)
42	Acad_Vernadsky
43	Admap (Compilation)
44	Marine model
45	WDMAM 2007 (grid8)
46	Compilation Getech
47	GRIMM_L120

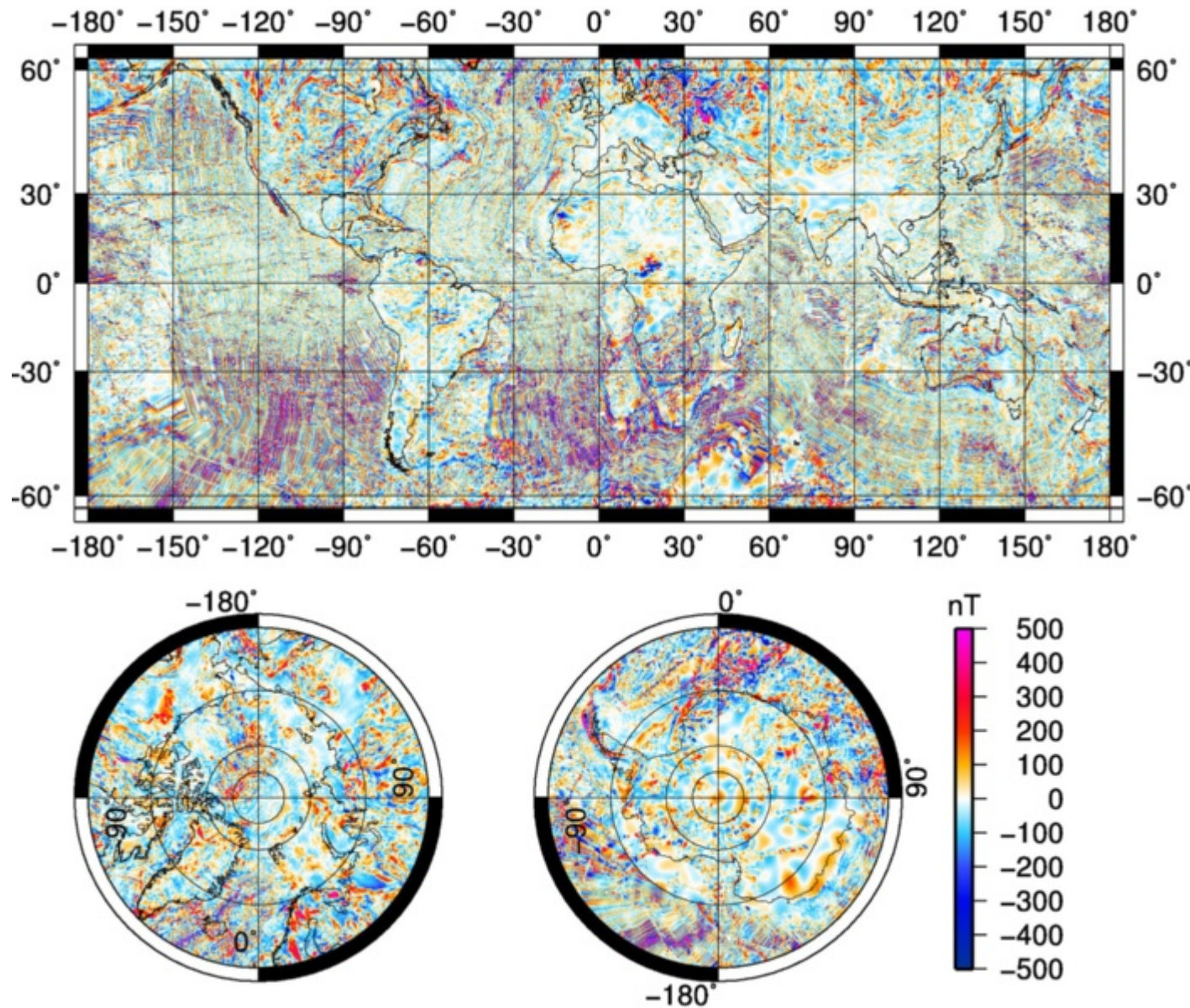
# WDMAM version 2.0

Lesur et al. (2016)





# WDMAM divided by main field strength/45,000 nT



**Continents** - mainly induced magnetization

Large anomalies  
Siberia, North America,  
Australia, South Africa

Smaller anomalies  
Western Europe, China,  
Central to north Africa,  
South America  
(partly due to lack of  
detailed data)

Contrast at the  
Tornquist-Teisseyre line

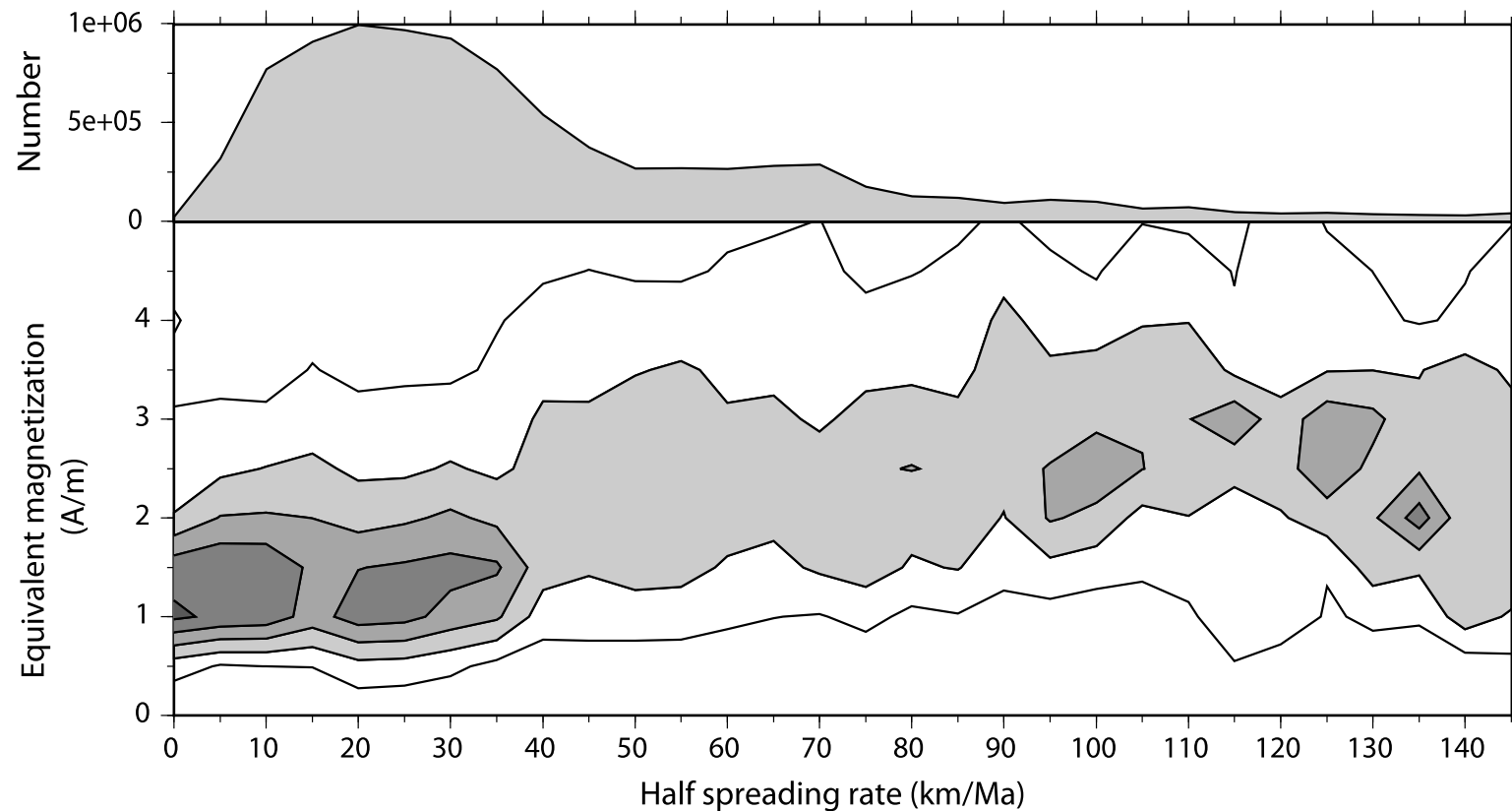
After Lesur et al. (2016)

# Oceanic area model - Equivalent magnetization

after Dyment et al. (2015)

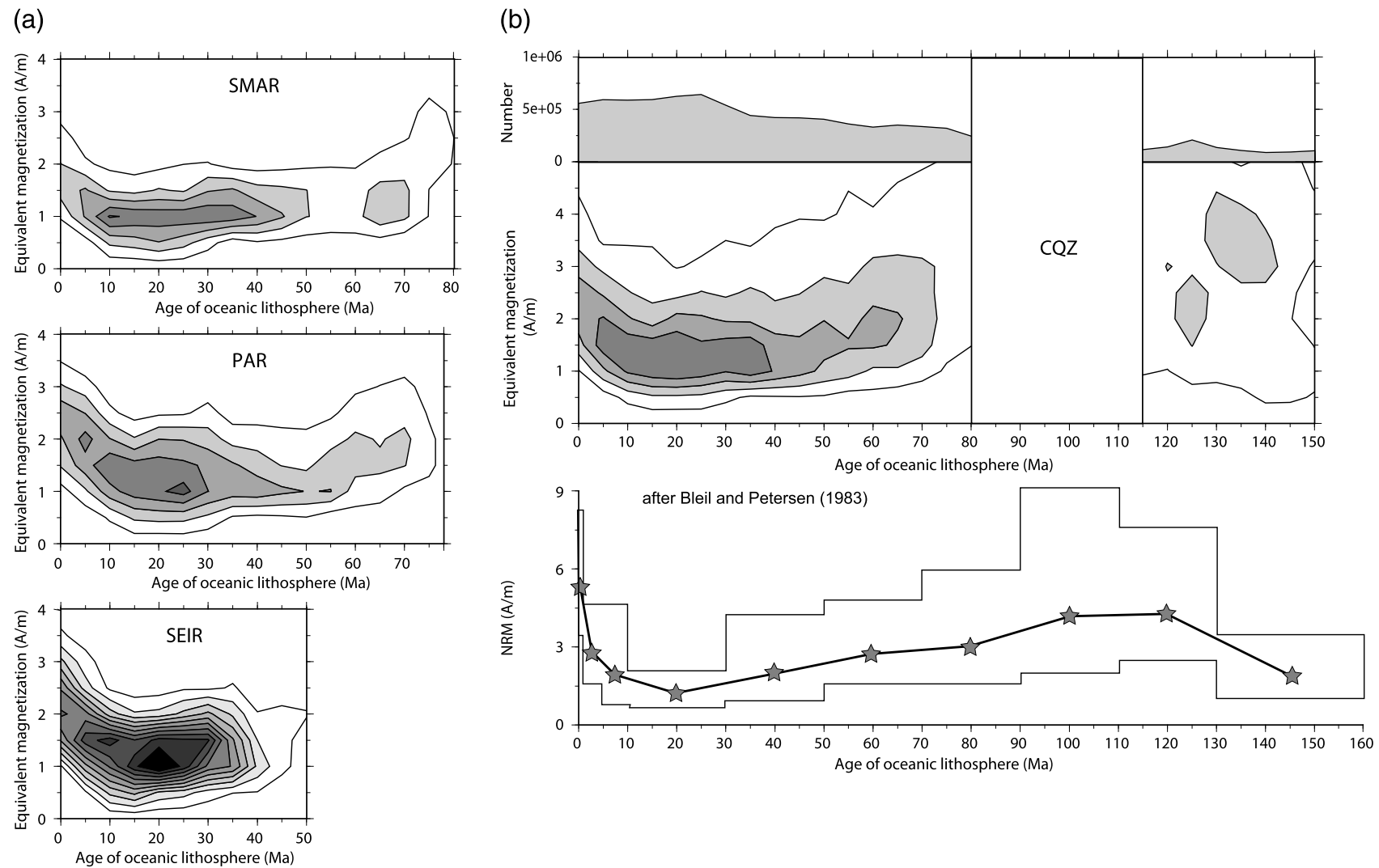
- (1) about twice stronger for areas formed at fast spreading rate
- (2) decreases significantly during the first 10 to 15 Ma, and slowly increases from 20 to 70 Ma.

## Equivalent magnetization versus spreading rate

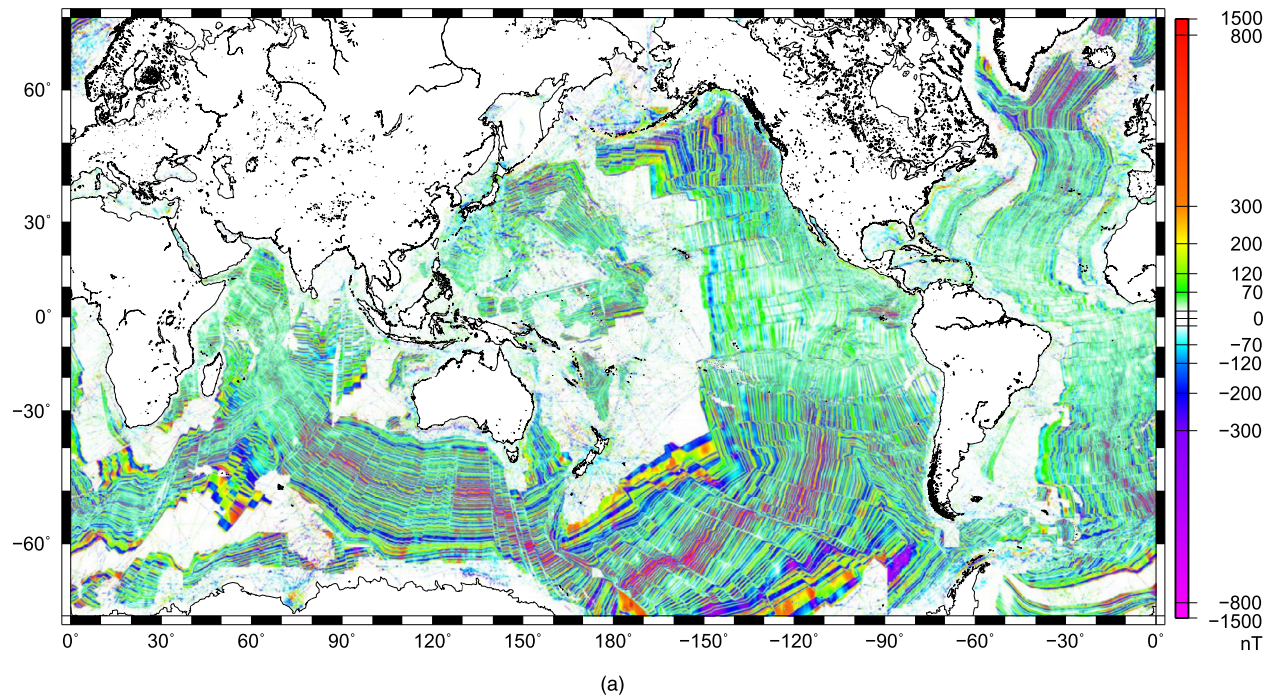


# Equivalent magnetization versus age

after Dyment et al. (2015)

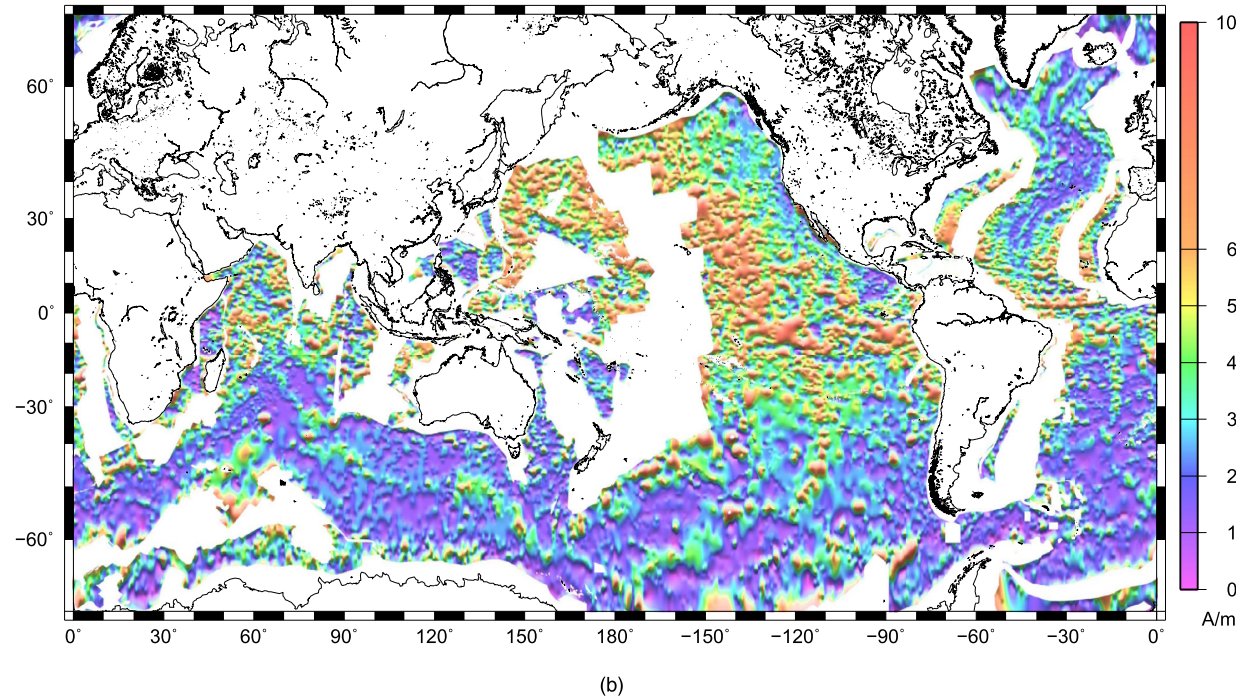


**Fig. 3.** Equivalent magnetization (assuming a 1 km-thick magnetized source layer mimicking the top of the basement) versus age of the oceanic lithosphere. See text for details on plot constructions. a) For selected areas, not normalized, from top to bottom: the southern Mid Atlantic Ridge and flanks (SMAR, contours every 6000 samples), the Pacific–Antarctic Ridge and flanks (PAR, contours every 8000 samples), and the Southeast Indian Ridge (SEIR, contours every 8000 samples). b) For all dated oceanic areas, normalized for each age to the number of measurements (top), contoured every multiples of 3% (medium), compared to the Natural Remanent Magnetization (NRM) measured on drilled samples (bottom; after [Bleil and Petersen, 1983](#): stars and thick line, average values; thin lines, uncertainties).



Oceans – mainly  
remanent magnetization

Marine magnetic anomaly model  
Magnetization vector  
(10 A/m, 1km thick layer  
beneath top basement assumed)  
from ocean floor age,  
geomagnetic polarity time scale  
& finite rotation for each plate



Adjust equivalent magnetization  
to fit the amplitude variations of  
model and real data on sliding  
windows of 100 to 400 km  
along each track line

After Dyment et al. (2015)

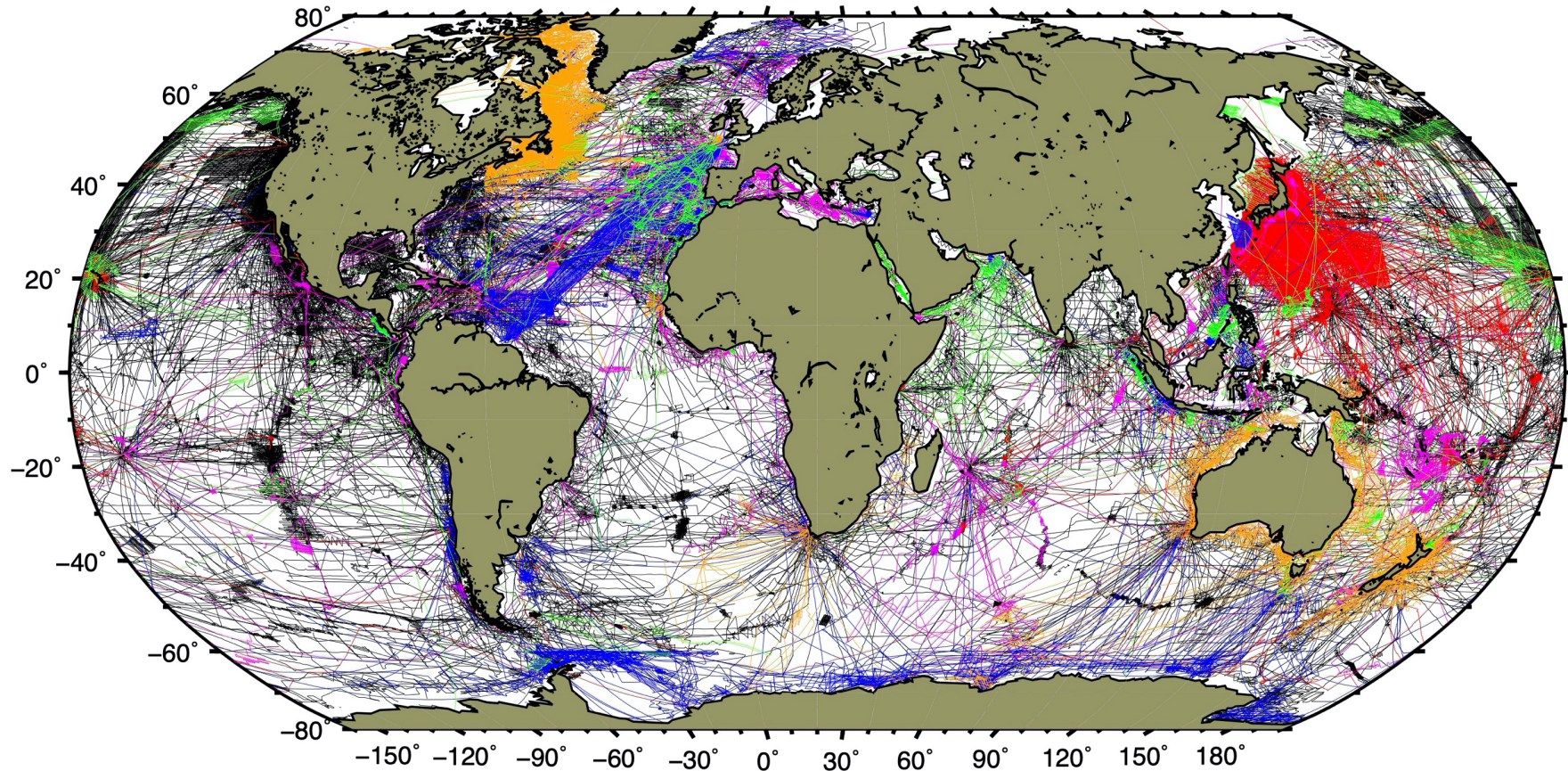
Fig. 2. Final products of our analysis as defined in Fig. 1. (a) The World Digital Magnetic Anomaly Map version 2 at sea level and (b) the equivalent magnetization (assuming a 1 km-thick magnetized source layer mimicking the top of the basement) over oceanic areas.

# Compilation of global marine magnetic anomaly data

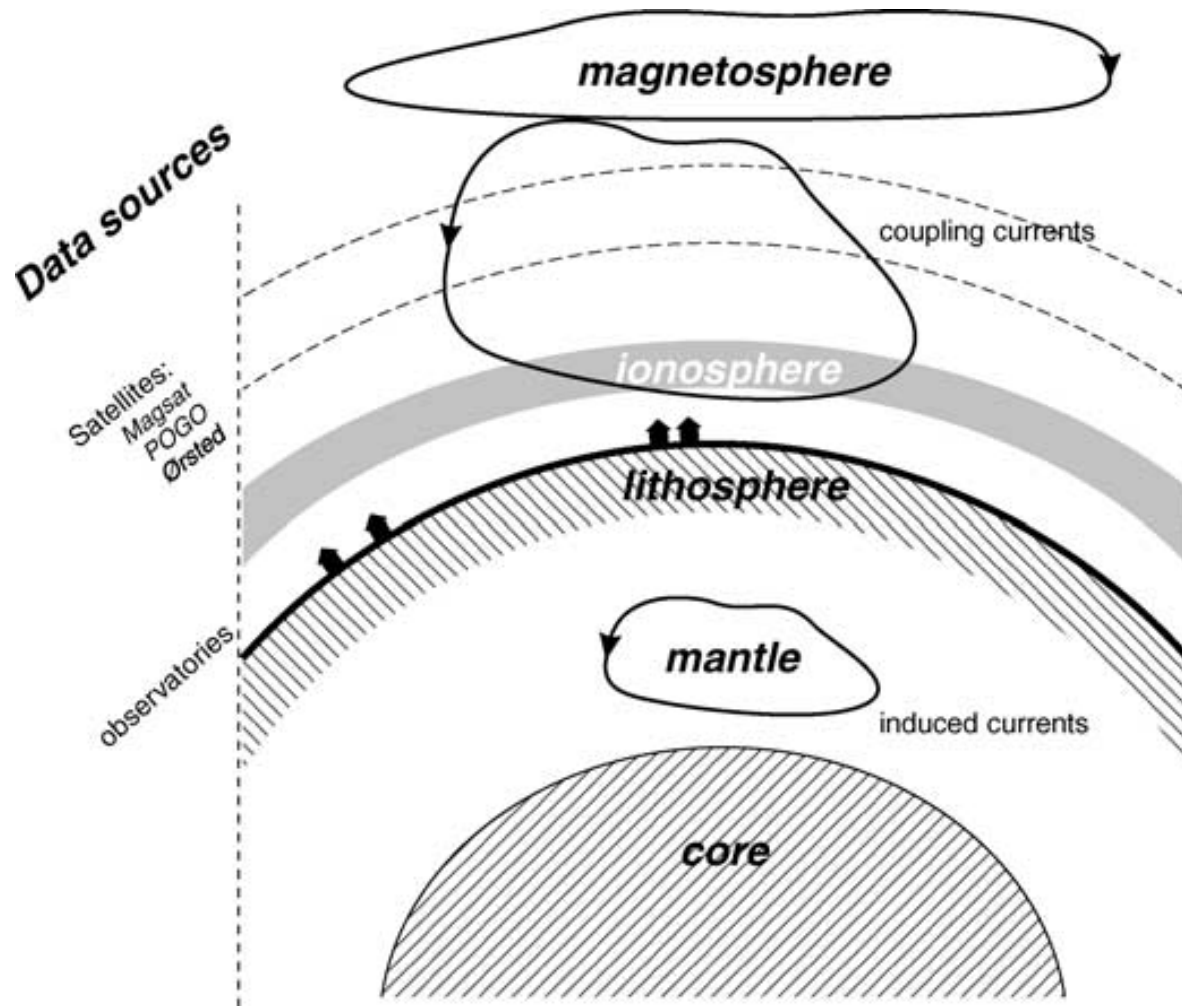
# Collected marine data

	Cruises	Records
GEODAS 5.0.10 (Quesnel et al., 2009)	2245	19072440
Global2019	3453	36679884

	Cruises	Records		Cruises	Records		Cruises	Records
— US	2102	18164546	— Japan	485	4605716	— ADMAP I	85	1651670
— Canada	93	2787816	— France	277	3822592	— Netherlands	69	816390
— Australia	69	1927098	— UK	99	1059249	— Spain	11	118938
— New Zealand	100	759532	— Germany	31	829982	— China	3	8798
— South Africa	17	58475	— Russia	10	53492	— Israel	2	15590



# 3 major components of geomagnetic field



(after Sabaka et al., 2002)

## Main field

Self-excited dynamo  
in the liquid outer core

## External field

Currents in the ionosphere  
and magnetosphere

## Anomaly

Lithospheric Magnetization

# Sources and variations

## of the 3 components of geomagnetic field

Component	Source	Spatial Variation	Temporal Variation
Main field	Outer core	Long wavelength Spherical harmonic degree $\sim < 15$	Secular variation ( $\sim > 1$ year)
External fields	Ionosphere & Magnetosphere	Long wavelength	1 hour to 1 year
Anomalies	Lithosphere	Short wavelength Spherical harmonic degree $\sim > 16$	No variation



# New leveling method

Ishihara (EPS, 2015) modified

to correct temporal variation due to external field

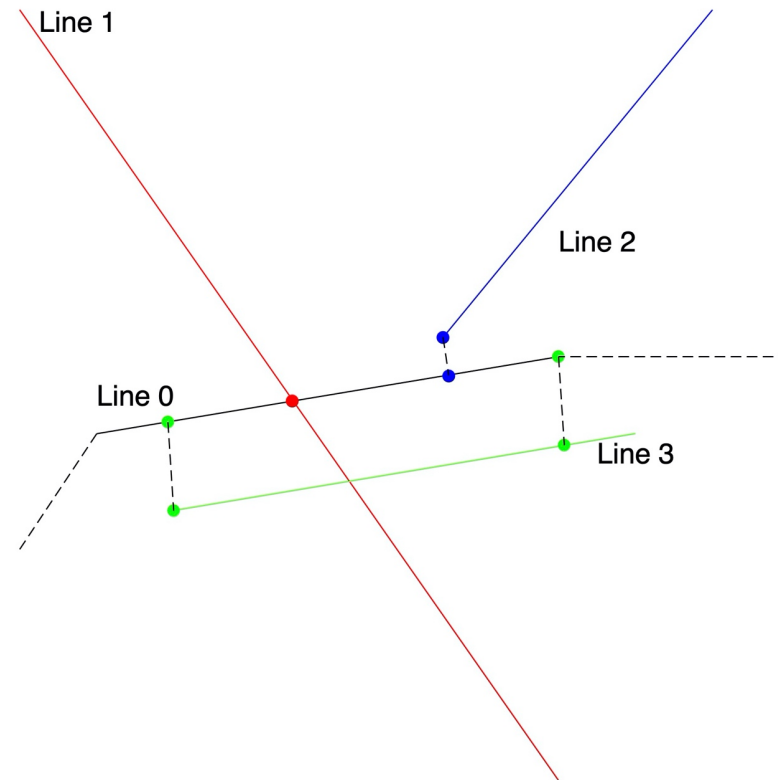
Ship's track is divided into straight lines. Crossover point or nearest points between each line pair is considered. Combination of weighted differences from neighboring line values and low-pass filtering. Weight of the anomaly difference between the nearest points is given as function of their distance, and the leveling correction is obtained by low-pass filtering of the weighted anomaly differences along the ship's track.

$$c_i = \frac{\sum_k G(t_k - t_i) \sum_j W(d_{jk}) (c_j + a_k - a_j)}{\{\sum_k G(t_k - t_i) \sum_j W(d_{jk})\}}$$

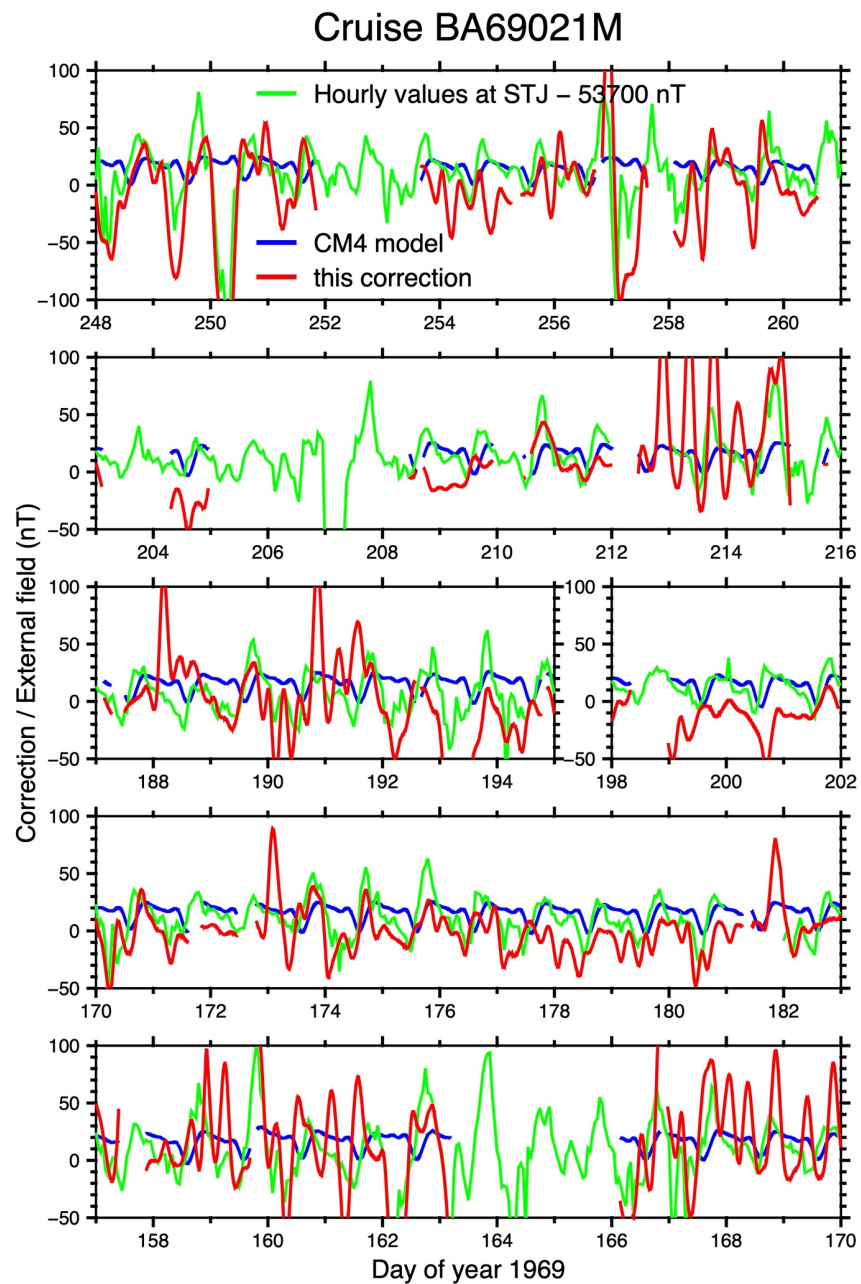
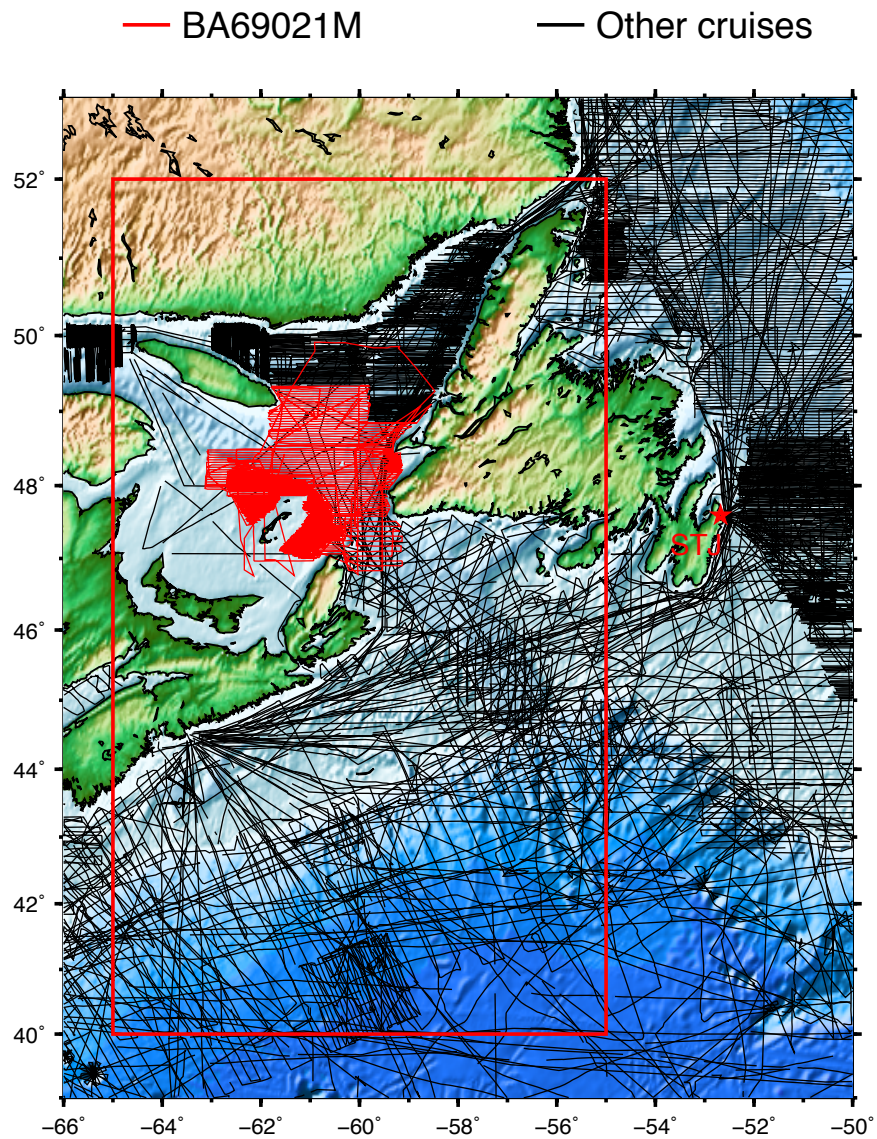
$W(d_{jk})$ : weight function of distance  
between j-th and k-th points

$G(t_k - t_i)$ : Gaussian filtering function  
of time difference between  
k-th and i-th points

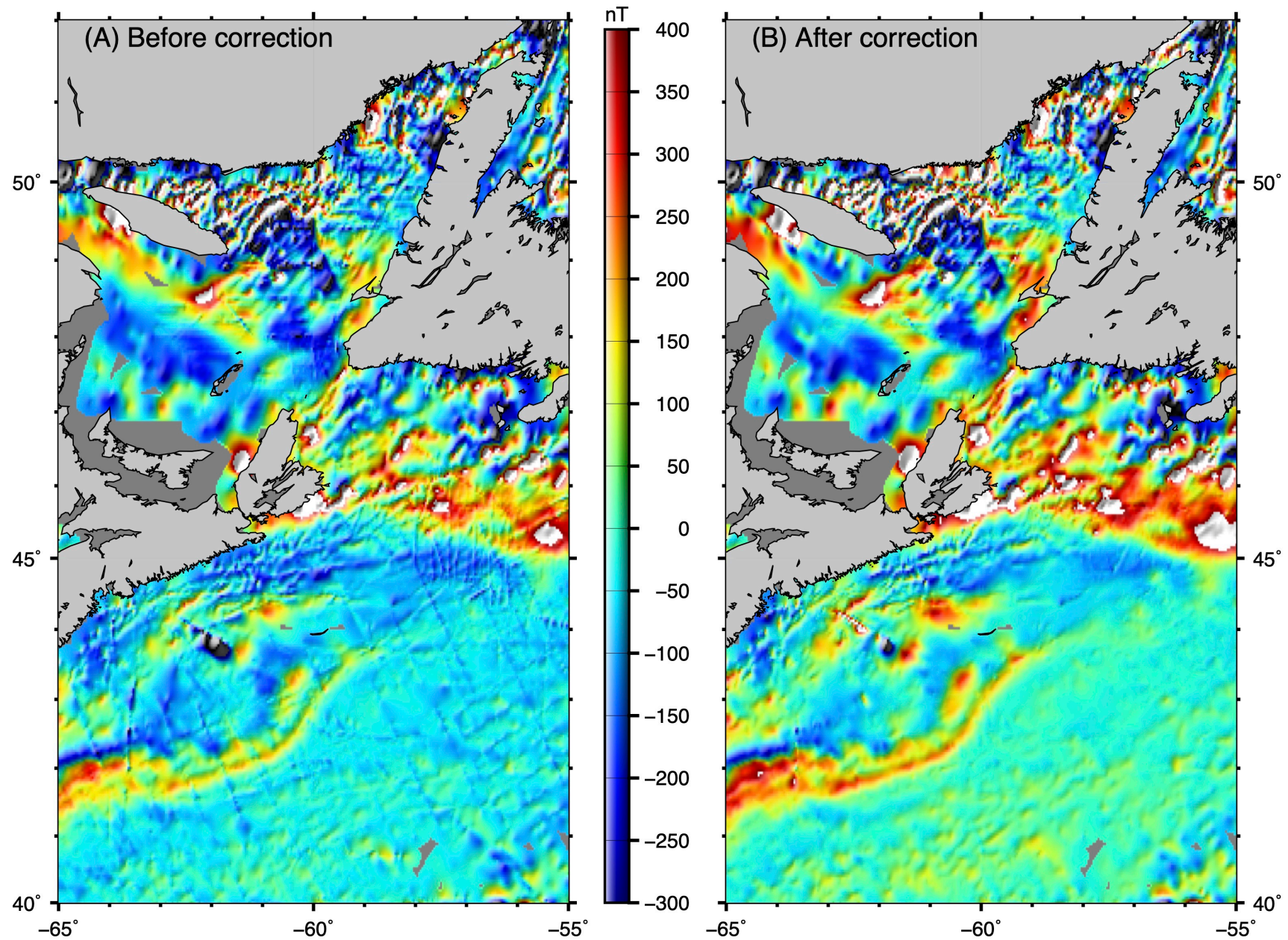
Correction  $c_i$  is obtained by  
iterated calculation.



# Gulf of St. Lawrence

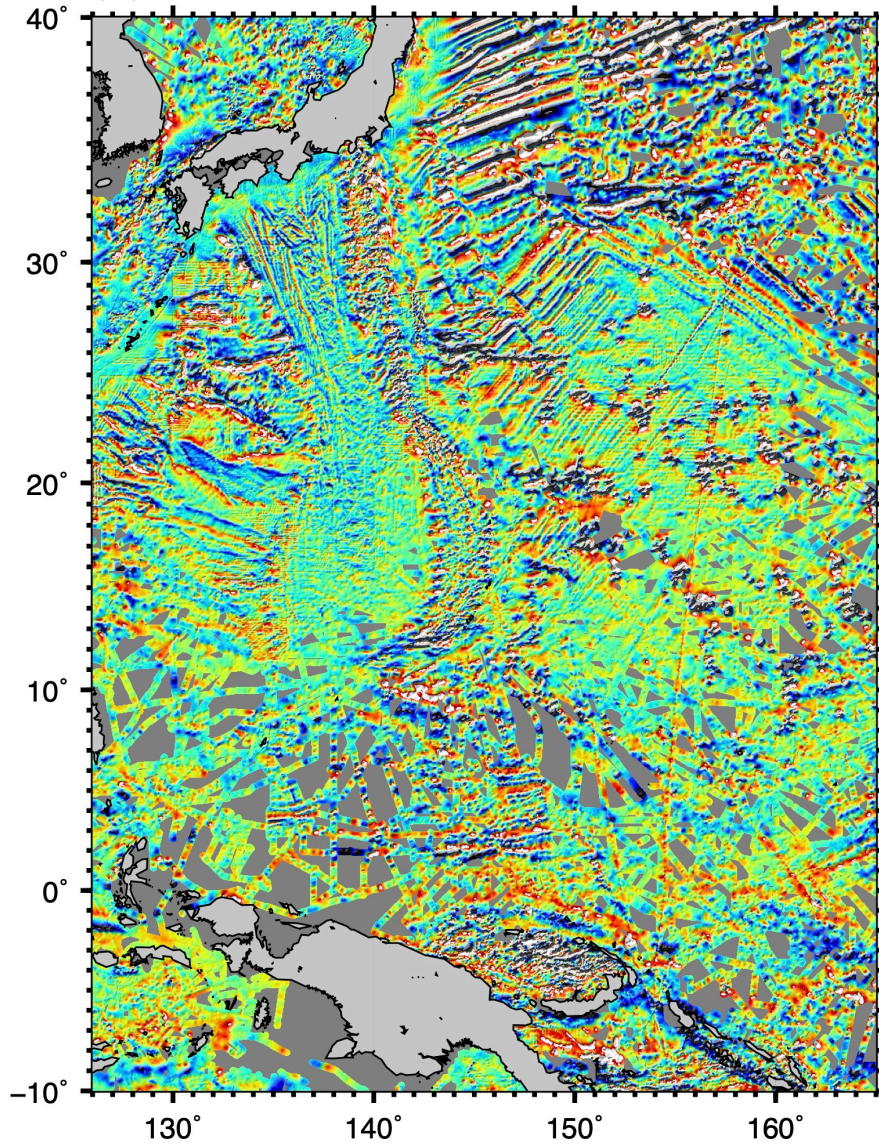


# Gulf of St. Lawrence, NW Atlantic Ocean

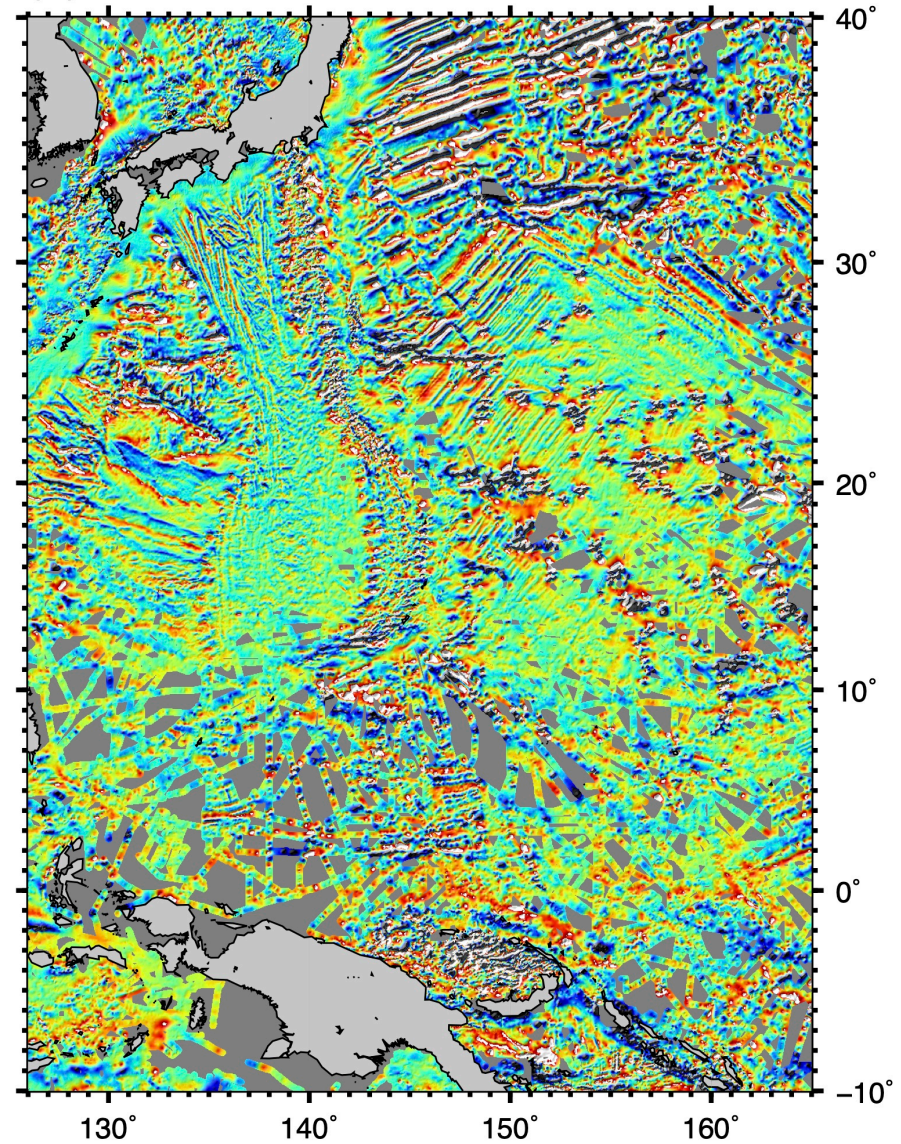


# Philippine Sea and NW Pacific Ocean

(A) Before correction



(B) After correction

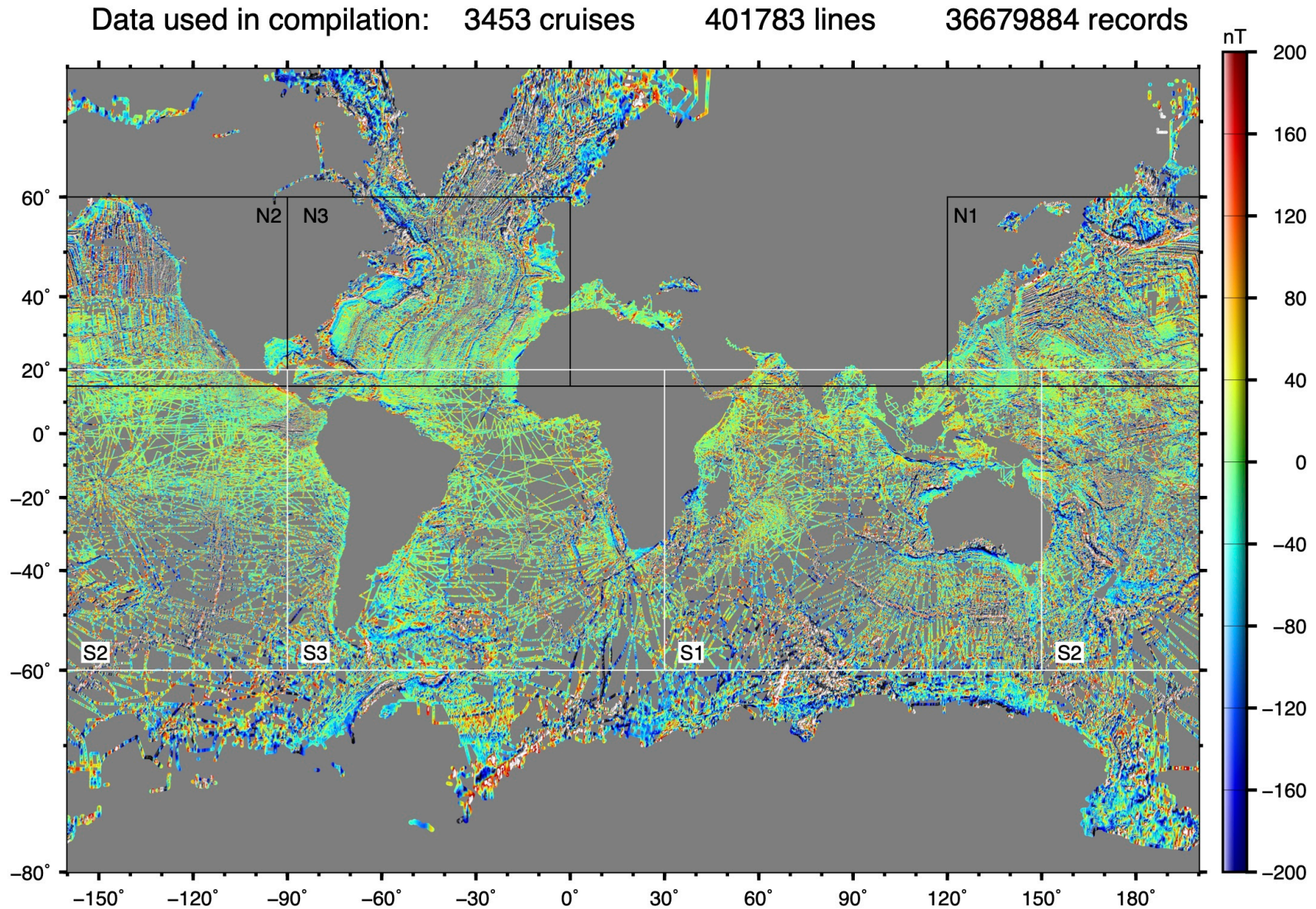


# Global marine magnetic anomaly map after leveling correction

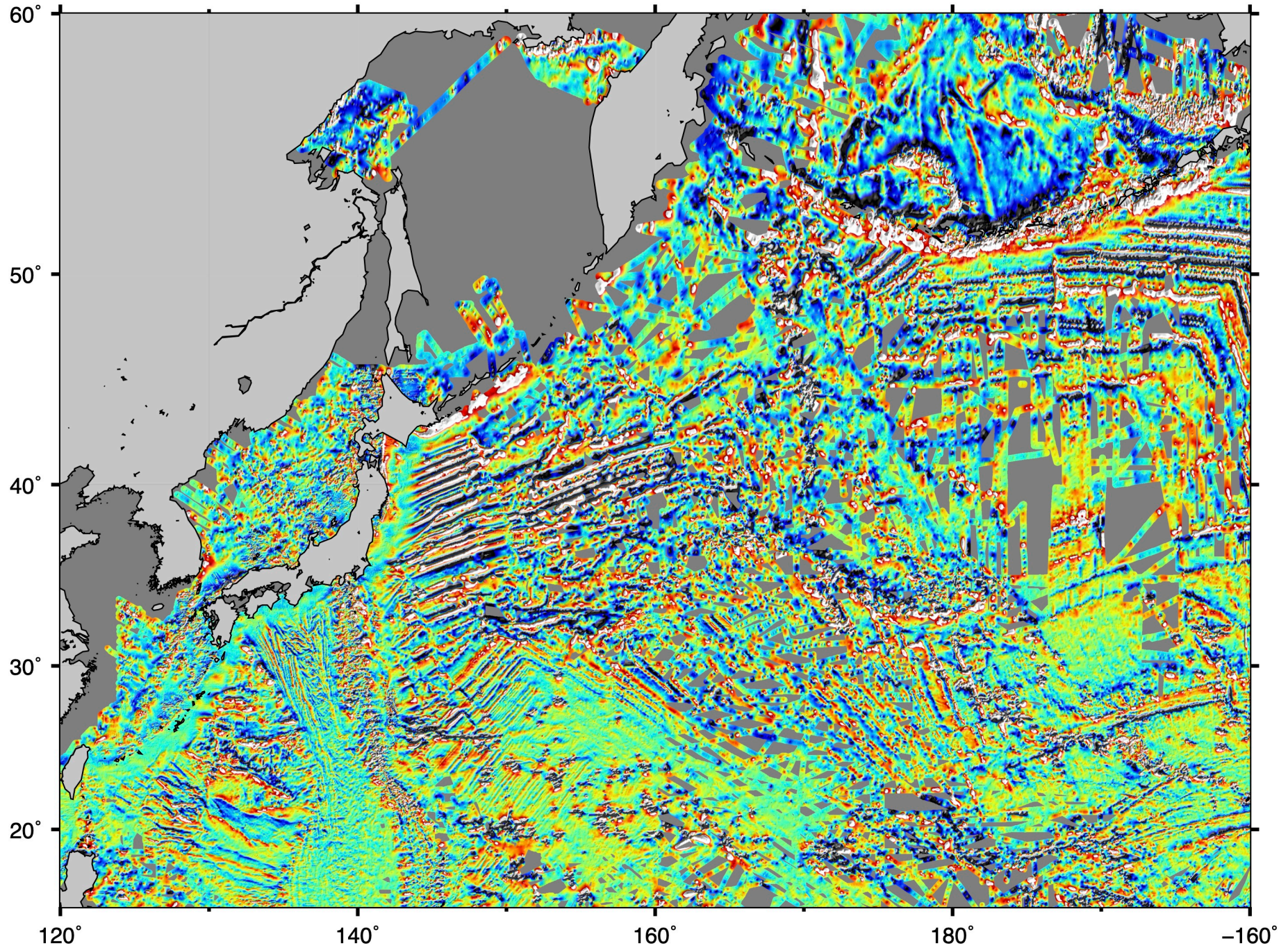
rms COD: 79.5 nT (before correction)

45.6 nT (after correction)

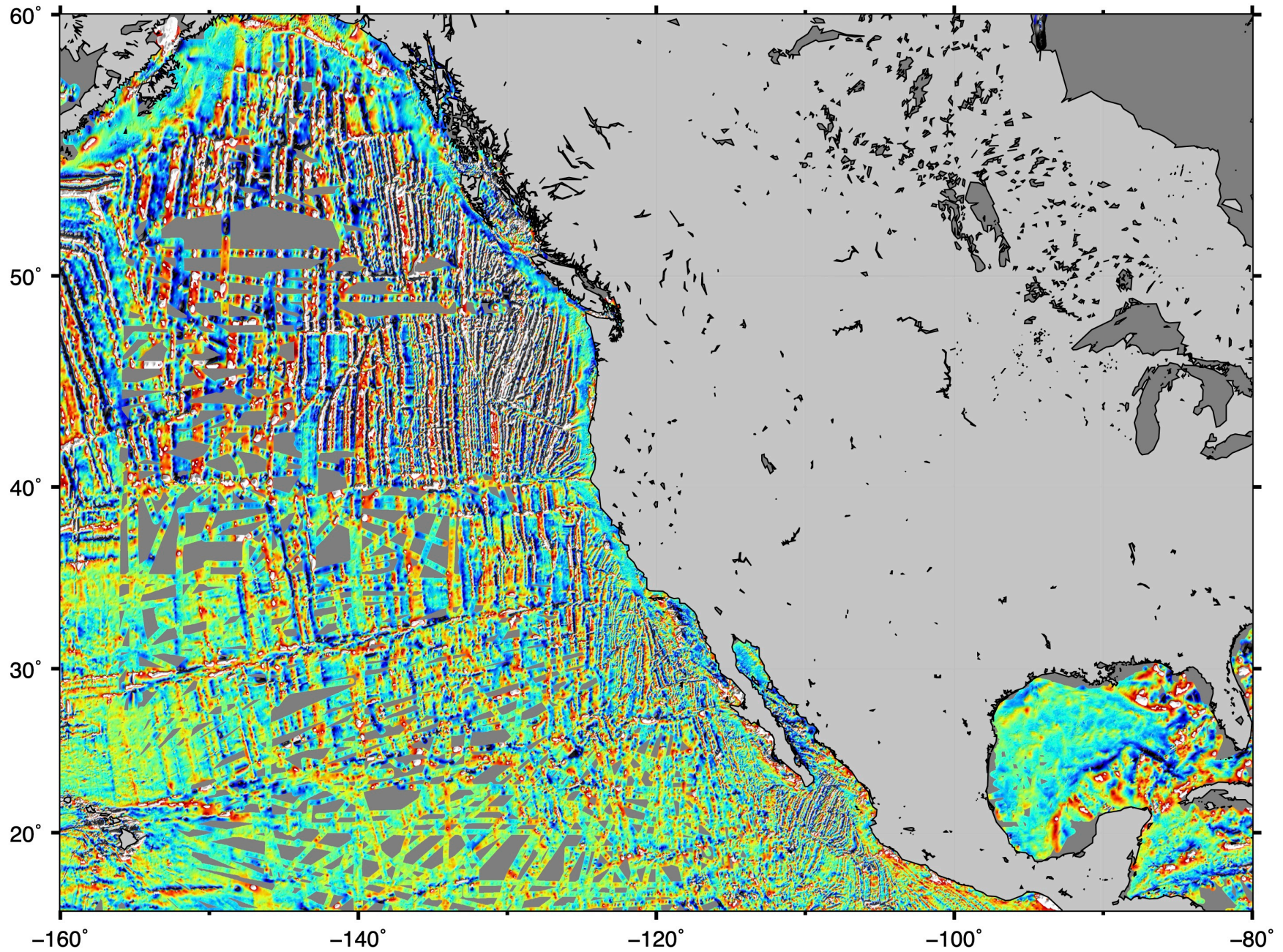
Total NO of CODs = 806466



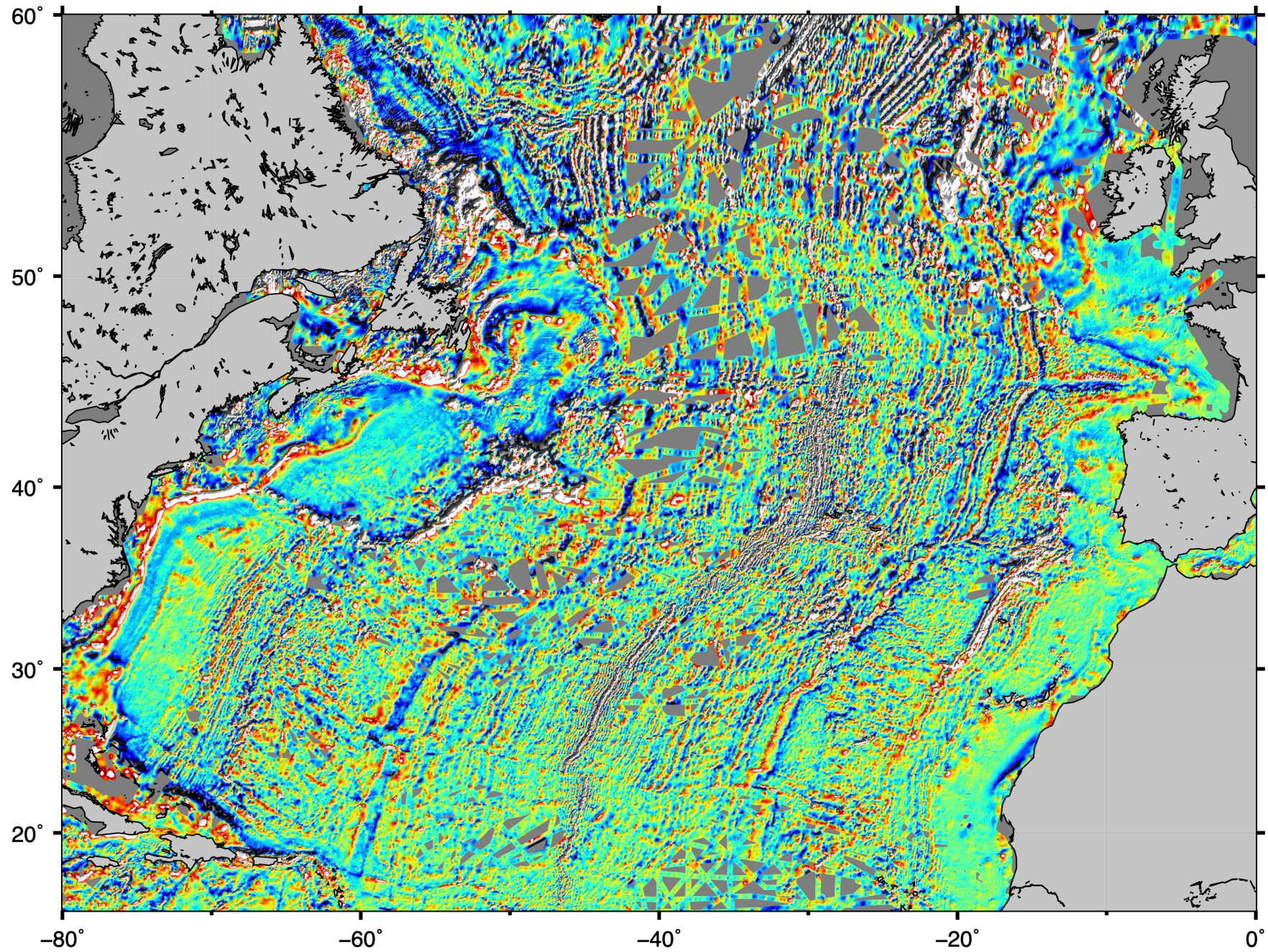
# N1: Northwest Pacific Ocean



# N2: Northeast Pacific Ocean

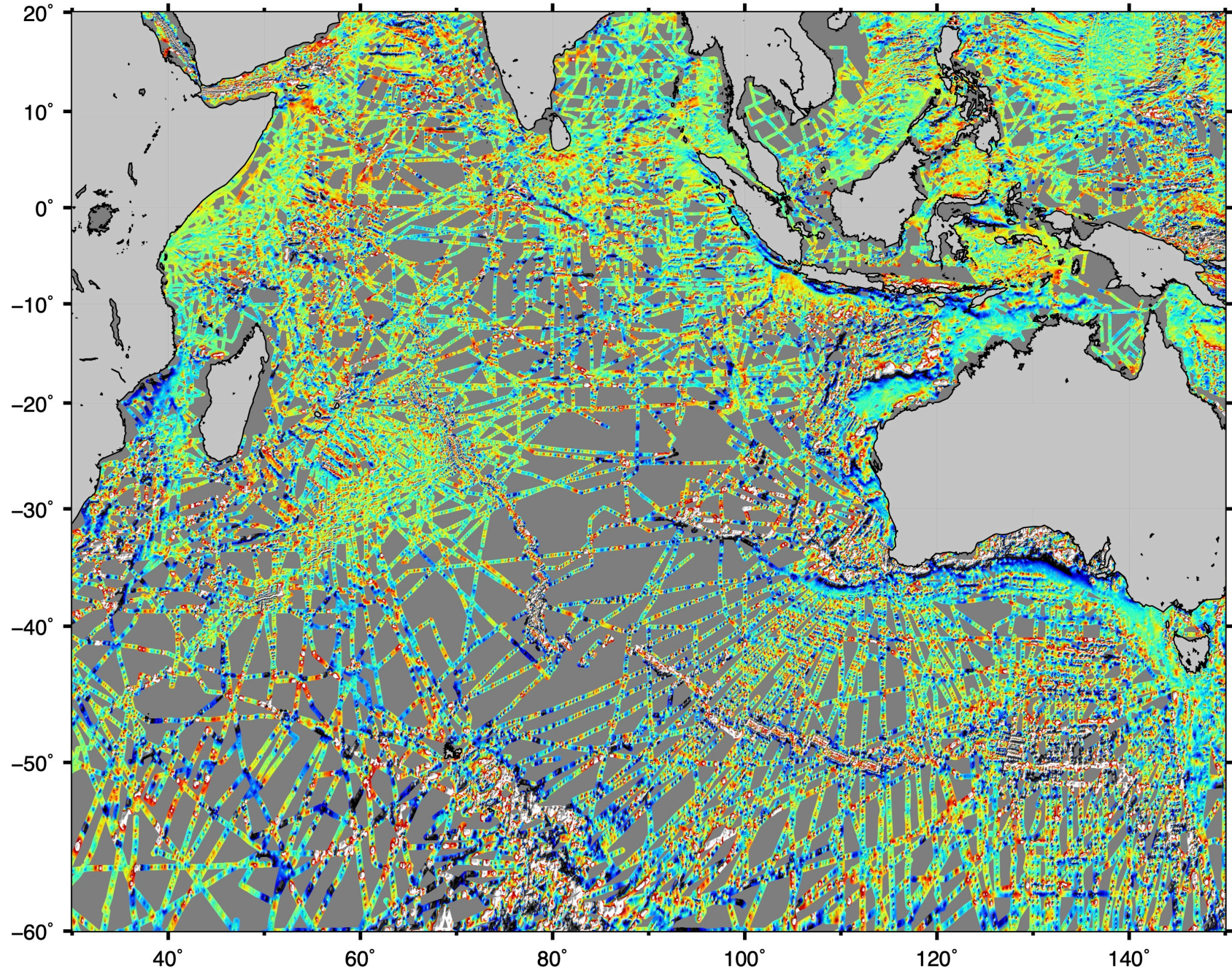


# N3: North Atlantic Ocean

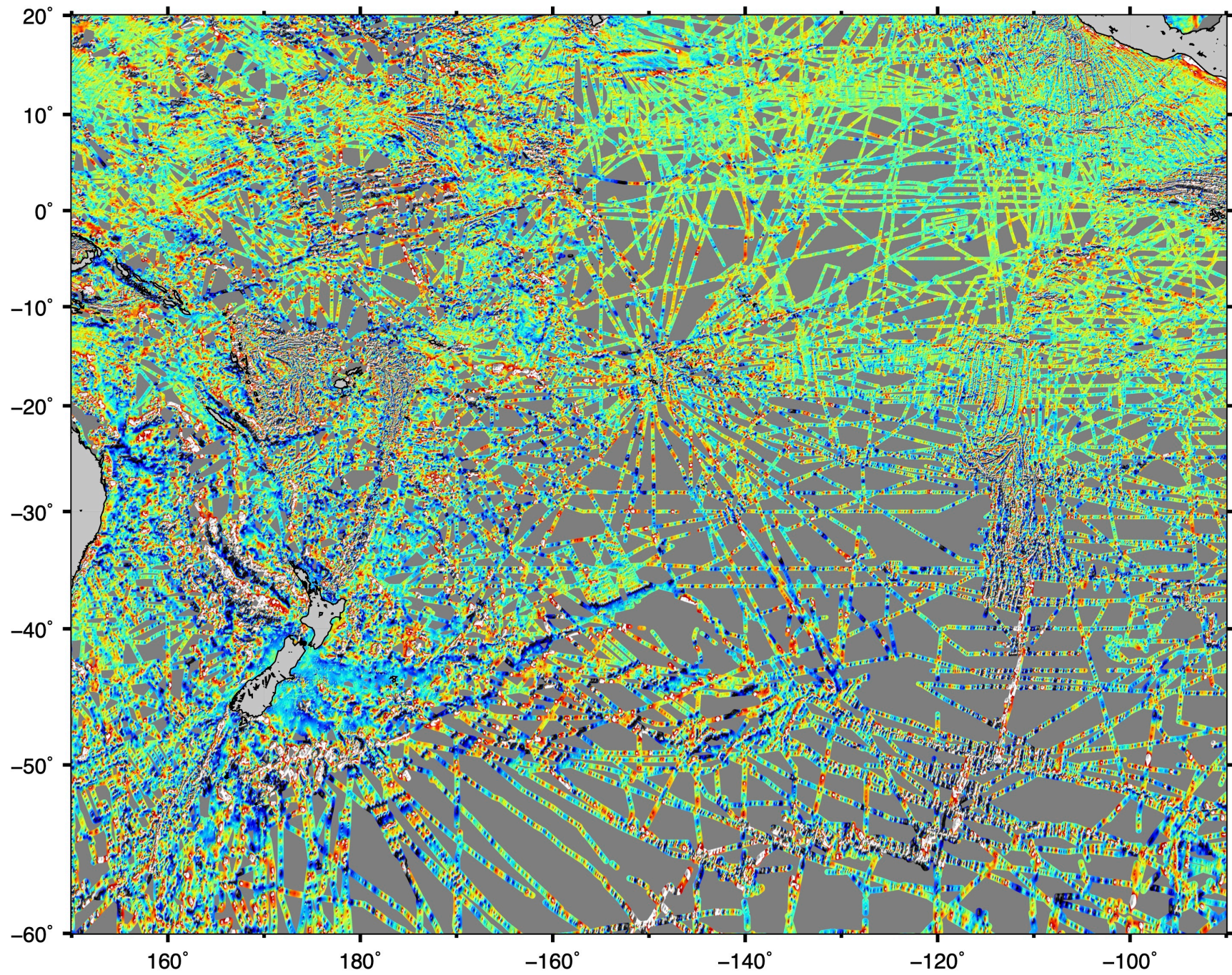




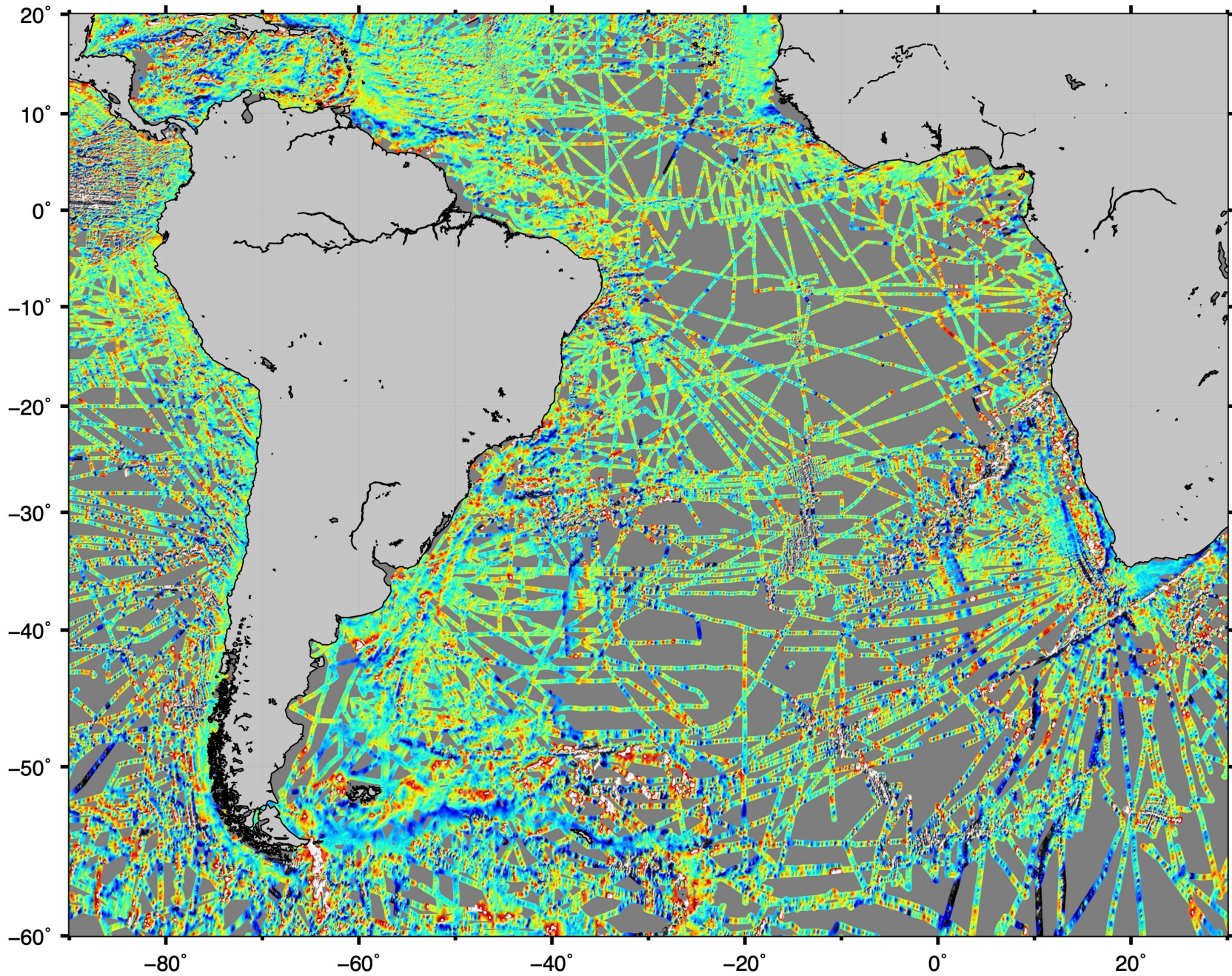
# S1: Indian Ocean



# S2: South Pacific Ocean



# S3: South Atlantic Ocean



# References

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- Ishihara, T. (2015) A new leveling method without the direct use of crossover data and its application in marine magnetic surveys: weighted spatial averaging and temporal filtering. *Earth, Planets and Space*, 67-11 DOI 10.1186/s40623-015-0181-7.
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