

# **A Case Study of on a Numerically Simulated Ionospheric Convection with a Global MHD Simulation**

Satoko Saita<sup>1</sup>, Shigeru Fujita<sup>2</sup>, Akira Kadokura<sup>3</sup>, Takashi Tanaka<sup>3</sup>, Yoshimasa Tanaka<sup>3</sup>

<sup>1</sup>*National Institute of Technology, Kitakyushu College*

<sup>1</sup>*Meteorological College*

<sup>1</sup>*Research Organization of Information and Systems, National Institute of Polar Research*

<sup>2</sup>*International Center for Space Weather Science and Education, Kyushu University*

A case study was conducted for simulated processes of magnetosphere-ionosphere (M-I) coupling process using a magnetohydrodynamic (MHD) simulation code developed by Tanaka (2010).

We combined the MHD simulation and solar wind parameters derived from the ACE satellite, and compared the ionospheric  $\mathbf{E} \times \mathbf{B}$  plasma drift obtained from the global MHD simulation and that obtained from the SuperDARN HF Radar Network.

The simulated plasma drift was not always reproducible under a southward interplanetary magnetic field (IMF) condition. We believe that the M-I boundary conditions in the global MHD simulation includes insufficient factors for the M-I coupling process.

## **References**

Tanaka T, A. Nakamizo, A. Yoshikawa, S. Fujita, H. Shinagawa, H. Shimazu, T. Kikuchi, and K. Hashimoto, Substorm convection and current system deduced from the global simulation, *J. of Geophys. Res.*, vol. 115 A05220, 2010.