

APPLICATION OF MULTIVARIATE STATISTICAL ANALYSIS TO CLASSIFICATION OF ANTARCTIC STONY METEORITES

Masaki EJIRI, Haruo SAKURAI, Minoru FUNAKI and Takesi NAGATA

National Institute of Polar Research, 9-10, Kaga 1-chome, Itabashi-ku, Tokyo 173

Abstract: Two standard multivariate statistical computer analyses, *i.e.* cluster and stepwise discriminant analyses, are applied to classification of antarctic stony meteorites of 1 E chondrite, 11 H chondrites, 14 L chondrites, 4 LL chondrites, 8 C chondrites, 15 achondrites and 1 mesosiderite. Six cases by changing combinations of several variables of magnetic, chemical and petrographical components give basically the same classification of sample grouping, except for a few samples. Also it is worthy of special mention that only two parameters of I_s and $I_s(\alpha)/I_s$ make a good classification of sample grouping.

1. Introduction

Magnetic characteristics of antarctic stony meteorites have been intensively studied by NAGATA (1979a, b), taking account of mineralogical and petrographical properties by examining a Urey-Craig-Mason diagram (UREY and CRAIG, 1953; MASON, 1962) and a Prior's rule (PRIOR, 1920) for chondrites, and he showed that stony meteorite groups can be satisfactorily identified on the basis of magnetic data alone.

Adding 16 samples to the above Nagata's study, two kinds of multivariate statistical analyses (*e.g.* BMDP, 1979) are applied to classification of antarctic stony meteorites; one is a cluster analysis which gives a grouping on the basis of n -dimensional Euclid distances of variables, and the other is a stepwise discriminant analysis based on the distribution of probabilities of samples belonging to each group. Effects of combination of variables on the classification are examined for 6 cases.

The main objective of this study is to clarify the differences of classification by magnetic characteristics, chemical and petrographical properties, and their combinations. Although two methods for 6 cases give basically the same classification, a few differences are found, and it has been proved that this kind of computer analysis is useful for classification and identification of magnetic, chemical and petrographical properties of antarctic stony meteorites.

2. Data of Antarctic Stony Meteorites

42 samples of antarctic stony meteorites have been used in Nagata's previous study (1979 a, b), and 16 samples are newly added in the present analysis, as shown in Table 1. Since a preliminary analysis indicates that 4 irons make a definite

Table 1. Samples of antarctic stony meteorites with their characteristics. Abbreviations are explained in the text.

| ID | NO. | | Sample ID | I1 | I2 | I3 | I4 | I5 | I6 | I7 | I8 | I9 | I10 | I11 | I12 | I13 | I14 |
|----|-----|----|-----------------|------|----|----|----|----|------|-----|-------|------|-------|-------|-------|-------|------|
| E | 1 | 1 | 1 Y-691 (a) | 48.0 | 97 | 0 | 3 | 0 | 12 | 764 | 22.18 | 1.86 | 0.089 | 11.47 | 0.48 | 11.92 | 5.6 |
| H | 2 | 1 | 2 Y-694 (d) | 32.3 | 94 | 6 | 0 | 0 | 23 | 685 | 12.45 | 1.52 | 0.081 | 5.32 | 12.32 | 8.19 | 6.8 |
| | | 2 | 3 Y-7301 (j) | 15.5 | 85 | 10 | 5 | 0 | | 660 | 7.21 | 0.77 | 0.05 | 5.05 | 18.92 | 9.36 | 10.9 |
| | | 3 | 4 Y-74371 | 33.5 | 95 | 5 | 0 | 0 | 10 | 635 | | | | | | | 7.7 |
| | | 4 | 5 Y-74647 | 27.9 | 94 | 6 | 0 | 0 | 14 | 659 | | | | | | | 7.1 |
| | | 5 | 6 Kesen | 34.4 | 95 | 5 | 0 | 0 | 8 | 670 | | | | | | | 6.5 |
| | | 6 | 7 Yonozu | 24.2 | 87 | 13 | 0 | 0 | 42 | 654 | | | | | | | 10.2 |
| | | 7 | 8 Seminole | 24.3 | 94 | 6 | 0 | 0 | 18 | 627 | | | | | | | 8.2 |
| | | 8 | 9 Mt. Baldr b | 27.3 | 88 | 10 | 2 | 0 | 10 | 650 | | | | | | | 9.5 |
| | | 9 | 10 Mt. Brown | 40.0 | 90 | 5 | 5 | 0 | | 640 | | | | | | | 9.9 |
| | | 10 | 11 Y-74054 | 24.0 | 98 | 2 | 0 | 0 | 11.6 | | | | | | | | |
| | | 11 | 12 Y-74115 | 23.9 | 96 | 4 | 0 | 0 | 25.5 | | | | | | | | |
| L | 3 | 1 | 13 Y-7305 (k) | 14.3 | 38 | 0 | 0 | 0 | | 624 | 7.64 | 0.96 | 0.06 | 7.57 | 13.10 | 7.96 | 10.3 |
| | | 2 | 14 Y-7304 (m) | 16.6 | 90 | 0 | 10 | 0 | 4 | 644 | 7.50 | 0.83 | 0.06 | 8.44 | 13.02 | 9.04 | 10.2 |
| | | 3 | 15 Y-74191 | 6.8 | 79 | 21 | 0 | 0 | 30 | 671 | 5.66 | 0.85 | 0.03 | 5.01 | 14.68 | 6.66 | 11.6 |
| | | 4 | 16 Y-74362 | 8.1 | 81 | 19 | 0 | 0 | 38 | 645 | | | | | | | 11.4 |
| | | 5 | 17 Fukutomi | 22.9 | 82 | 18 | 0 | 0 | 20 | 700 | 9.83 | 1.33 | 0.02 | 6.37 | 11.62 | 7.39 | 10.2 |
| | | 6 | 18 Mino | 11.0 | 80 | 20 | 0 | 0 | 3 | 653 | 7.86 | 1.16 | 0.05 | 5.88 | 14.48 | 6.78 | 12.1 |
| | | 7 | 19 ALH-769 | 8.4 | 65 | 35 | 0 | 0 | 160 | 680 | | | | | | | 15.7 |
| | | 8 | 20 Dalgety Down | 9.7 | 85 | 14 | 0 | 0 | 117 | 648 | | | | | | | 10.5 |
| | | 9 | 21 Bjurböle | 13.0 | 85 | 10 | 0 | 0 | | 660 | | | | | | | 10.8 |
| | | 10 | 22 Barratta | 12.0 | 80 | 15 | 3 | 0 | | 655 | | | | | | | 11.6 |
| | | 11 | 23 Homestead | 10.0 | 80 | 15 | 5 | 0 | | 650 | | | | | | | 12.3 |
| | | 12 | 24 Y-74190 | 9.3 | 83 | 6 | 11 | 0 | 5.8 | | | | | | | | |
| | | 13 | 25 Y-74354 | 21.8 | 93 | 7 | 0 | 0 | 65.5 | | | | | | | | |
| | | 14 | 26 ALH-77260 | 2.0 | 70 | 25 | 5 | 0 | 115 | | | | | | | | |
| LL | 4 | 1 | 27 Y-74442 | 6.0 | 45 | 35 | 20 | 0 | 85 | 680 | 2.48 | 0.99 | 0.015 | 4.84 | 17.89 | 2.51 | 22.6 |
| | | 2 | 28 Y-74646 | 3.2 | 19 | 7 | 74 | 0 | 20 | 720 | 1.96 | 1.01 | 0.03 | 4.59 | 19.02 | 1.94 | 31.0 |

Table 1. (Continued).

| ID | NO. | | Sample ID | I1 | I2 | I3 | I4 | I5 | I6 | I7 | I8 | I9 | I10 | I11 | I12 | I13 | I14 |
|----|-----|---|----------------|--------|-----|----|----|-----|-------|-----|------|-------|-------|------|-------|--------|------|
| LL | 4 | 3 | 29 St. Severin | 4.7 | 45 | 55 | 0 | 0 | 500 | 700 | | | | | | | 20.5 |
| | | 4 | 30 ALH-764 | 5.1 | 88 | 12 | 0 | 0 | 87.5 | | | | | | | | |
| C | 5 | 1 | 31 Orgueil | 11.9 | 0 | 0 | 0 | 100 | (3.4) | | | | | | | | |
| | | 2 | 32 Ivuna | 11.2 | 0 | 0 | 0 | 100 | (3.4) | | | | | | | | |
| | | 3 | 33 Y-693 (c) | 10.8 | 0 | 0 | 0 | 100 | 157 | 540 | 0.06 | 0.00 | 0.08 | 3.62 | 27.95 | 999.00 | 4.0 |
| | | 4 | 34 Y-74662 | 0.81 | 5 | 0 | 85 | 10 | 149 | 580 | 0.00 | 0.00 | 0.06 | 7.38 | 22.53 | | |
| | | 5 | 35 Leoville | 10.3 | 6 | 0 | 0 | 94 | 34 | 575 | | | | | | | 4.0 |
| | | 6 | 36 Allende | 0.61 | 0 | 0 | 95 | 5 | 143 | 576 | | | | | | | |
| | | 7 | 37 Karoonda | 7.8 | 0 | 0 | 0 | 100 | 155 | 548 | | | | | | | |
| | | 8 | 38 Makoia | 8.0 | 0 | 0 | 0 | 100 | (2) | 570 | 0.00 | 0.00 | 0.00 | 6.74 | 25.43 | | |
| AA | 6 | 1 | 39 ALH-78113 | 5.33 | 100 | 0 | 0 | 0 | 24 | | | | | | | | |
| AD | 7 | 1 | 40 Y-692(b) | 0.19 | 81 | 0 | 19 | 0 | 42 | 780 | 0.66 | 0.004 | 0.003 | 1.34 | 12.58 | 165.00 | 4.0 |
| | | 2 | 41 Y-74013 | 0.17 | 56 | 0 | 44 | 0 | 10 | 792 | | | | | | | 4.0 |
| | | 3 | 42 Y-74037 | 0.32 | 100 | 0 | 0 | 0 | | | | | | | | | |
| | | 4 | 43 Y-74097 | 0.32 | 100 | 0 | 0 | 0 | 13 | | | | | | | | |
| | | 5 | 44 Y-74648 | 0.20 | 100 | 0 | 0 | 0 | 85 | | | | | | | | |
| | | 6 | 45 Y-75032 | 0.042 | 100 | 0 | 0 | 0 | 93 | | | | | | | | |
| AE | 8 | 1 | 46 Y-74159 | 0.061 | 100 | 0 | 0 | 0 | 265 | | | | | | | | |
| | | 2 | 47 Y-74450 | 0.22 | 100 | 0 | 0 | 0 | 58 | | | | | | | | |
| | | 3 | 48 ALH-765 | 0.076 | 100 | 0 | 0 | 0 | 15 | | | | | | | | |
| | | 4 | 49 ALH-78040 | 0.83 | 98 | 2 | 0 | 0 | 90 | | | | | | | | |
| | | 5 | 50 ALH-77302 | 0.012 | 100 | 0 | 0 | 0 | 24 | | | | | | | | |
| AH | 9 | 1 | 51 Y-7308 (1) | 0.53 | 100 | 0 | 0 | 0 | 13 | 792 | 0.39 | 0.012 | 0.007 | 0.75 | 16.00 | 32.5 | 4.0 |
| AU | 10 | 1 | 52 ALH-74123 | 6.45 | 93 | 7 | 0 | 0 | 41 | | | | | | | | |
| | | 2 | 53 Y-74659 | 2.23 | 84 | 9 | 7 | 0 | 76 | | | | | | | | |
| M | 11 | 1 | 54 ALH-77219 | 45.00 | 92 | 8 | 0 | 0 | 46.3 | | | | | | | | |
| I | 12 | 1 | 55 ALH-77255 | 184.50 | 98 | 2 | 0 | 0 | 9 | | | | | | | | |
| | | 2 | 56 ALH-762 | 210.00 | 100 | 0 | 0 | 0 | 6 | | | | | | | | |
| | | 3 | 57 DRP-78003 | 199.00 | 100 | 0 | 0 | 0 | 7.5 | | | | | | | | |
| | | 4 | 58 DRP-78007 | 176.50 | 100 | 0 | 0 | 0 | 4 | | | | | | | | |

group well-separated from others, these are not adopted to this study. These samples are identified chemically and petrographically as 1 enstatite (E) chondrite, 11 olivine-bronzite (H) chondrites, 14 olivine-hypersthene (L) chondrites, 4 olivine-pigeonite (LL) chondrites, 8 carbonaceous (C) chondrites, 15 achondrites (A) sub-classified into 1 aubrite (AA), 6 diogenites (AD), 5 eucrites (AE), 1 howardite (AH), and 2 ur-eilites (AU), and 1 mesosiderite (M).

14 variables are defined as; I_s is a saturation magnetization (usually denoted by I_s), I_2 to I_5 are ratios of the saturation magnetization to α -phase (kamacite), $(\alpha+\gamma)$ -phase, γ -phase and the total saturation magnetization (conventionally denoted by $I_s(\alpha)/I_s$, $I_s(\alpha+\gamma)/I_s$, $I_s(\gamma)/I_s$, and $I_s(Mt)/I_s$, respectively), I_6 is a magnetic coercive force (H_c) as a representative of the structure-dependent magnetic parameter, and I_7 is a main magnetic transition temperature (θ_c) in the cooling branch of thermomagnetic curve, in the case of stony meteorite, representing Curie point of magnetite in C chondrite, Curie point of Ni-poor α -phase FeNi in E chondrite and some achondrites, and γ to α transition temperature in H and L chondrites and some other achondrites. I_8 to I_{14} are the weight per cent of abundances of metallic iron (Fe°), Ni° , Co° , FeS, FeO, and the ratios of Fe° to Ni° and Ni° to $\text{Fe}^\circ + \text{Ni}^\circ$, respectively. In the analysis, the samples whose adopted variable values are not available are omitted.

3. Computer Multivariate Statistical Analysis

Several methods of computer multivariate statistical analyses have been developed well recently. A cluster analysis is one of intuitively perceivable methods, which uses, as a measure for grouping, N (number of valuables)-dimensional distances, such as a p -root of summation of distance to the power of p ($p=2$ for a Euclid distance being used in this analysis), etc. The distance between adjacent samples or between a center of an adjacent group and a sample is called an amalgamation distance which is used for this classification.

On the other hand, a stepwise discriminant analysis calculates a probability P_{ij} that a sample j belongs to i group, and finds a maximum value of the distribution of P_{ij} ($i=1, 2, \dots, M$: present number of groups); if $P_{Ij} = \max(P_{ij})$, then a sample j is most likely to be a member of a group I (For definition of P_{ij} and analysis procedure, see BMDP, 1979). The result P_{ij} ($i=1, 2, \dots, M$, and $j=1, 2, \dots, N$, N : total number of samples) is an $M \times N$ matrix or presented by M -dimensional structure for each sample. However, this paper employs a two-dimensional presentation by introducing a new coordinate system with two canonical variables in the following way. The linear combinations of variables in each set that have maximum correlation are the first coordinates in the new systems. Then a second linear combination in each set is sought such that the correlation between these is the maximum of correlations between such combinations as are uncorrelated with the first linear combinations. The procedure is continued until the two new coordinate systems are

completely specified (*e.g.* ANDERSON, 1958).

Among many possible combinations of variables I1 to I14, six cases are chosen with a view to investigate the effects of magnetic characteristics, chemical and petrographical properties on classification of antarctic stony meteorites.

Case 1 (I1 to I5) treats pure magnetic components of variables, neglecting other quantities. Case 2 (I1 to I5, and I6) adds I6 (H_c) to case 1. Case 3 (I1 to I5, and I7) adds I7 (θ_c) to case 1. Case 4 (I1 and I2) is a special case to study how only two variables I1 (I_s) and I2 ($I_s(\alpha)/I_s$) act effectively on the classification. Case 5 (I8 to I14) studies chemical and petrographical components of variables. Finally, case 6 (I1 to I5 and I8 to I14) is a combination of case 1 and case 4, trying to use characteristics of the both. Comparison among all cases reveals not only roles of each component of classification, but also properties of each antarctic stony meteorite.

4. Results and Discussions

A total of 58 samples of antarctic stony meteorites are used in this study. As previously mentioned, a preliminary analysis indicates that 4 irons make one definite group well-separated from others, and so these are omitted in this result.

The results of a stepwise discriminant analysis (hereinafter abbreviated to STW) for 6 cases are illustrated in Figs. 1 to 6, respectively in the form of the canonical variable coordinate system. As for a cluster analysis (hereinafter abbreviated to CLT), one example of case 2 is depicted in the amalgamation distance vertical tree diagram (Fig. 7) and its shaded form (Fig. 8).

Note that in the following a suffix indicates an identification number within a group, *e.g.* L₁₄ means a 14th sample of a group L listed in Table 1.

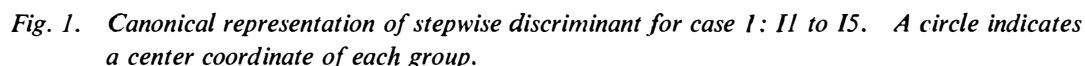
4.1. Case 1: I1 to I5 (Fig. 1)

One E chondrite Y-691 (a) and one mesosiderite ALH-77219 make one group by themselves. This is true for all cases and analyses.

H₂ Y-7301 (j) belongs to L-chondrite group and is near L₂ Y-7304 (m), L₉ Bjurböle, L₁₀ Barratta and L₆ Mino by STW or nearest to L₉ by CLT. H₉ Mt. Brown is placed between H and M groups.

L₁ Y-7305 (k) belongs to LL-chondrite group by STW but it is in-between an achondrite, L- and LL-chondrites groups and has long amalgamation distances to all groups by CLT. L₅ Fukutomi belongs to H-chondrite group and is near H₆ Yonozu but show a long amalgamation distance by CLT. L₁₃ Y-74354 belongs to H-chondrite group and is near H₁₀ and H₁₁ Homestead by STW and near H₇ Allan Hills-769 by CLT. L₇ Allan Hills-769 and L₁₄ ALH-77260 are in-between L- and LL-chondrites groups, and have long amalgamation distances to both groups.

LL₂ Y-74646 belongs to C-chondrite group but is in-between LL-chondrite and C-chondrite groups by STW, and CLT indicates the long amalgamation distance to



4.2. Case 2: I2 to I5, and I6 (Figs. 2, 7 and 8)

L₅ Fukutomi and L₁₃ Y-74354 belong to H-chondrite group. L₅ is near H₆ and L₁₃ is near H₁₀ and H₁₁ by both STW and CLT.

LL₂ approaches to C-chondrite group compared with case 1, but still has a long amalgamation distance. LL₄ is now belonging to the achondrite group apart from LL-chondrite group.

H₂ Y-7301 (j) belongs again to L-chondrite group and is near to L₉ Bjarboele by both STW and CLT. H₉ Mt. Brown is existing in-between H and M groups.

STW shows L₅ Fukutomi belongs to H-chondrite group and is near H₆ Mino, but L₅ belongs to its own L-chondrite group by CLT. L₁ Y-7305 (k) and L₇ Allan

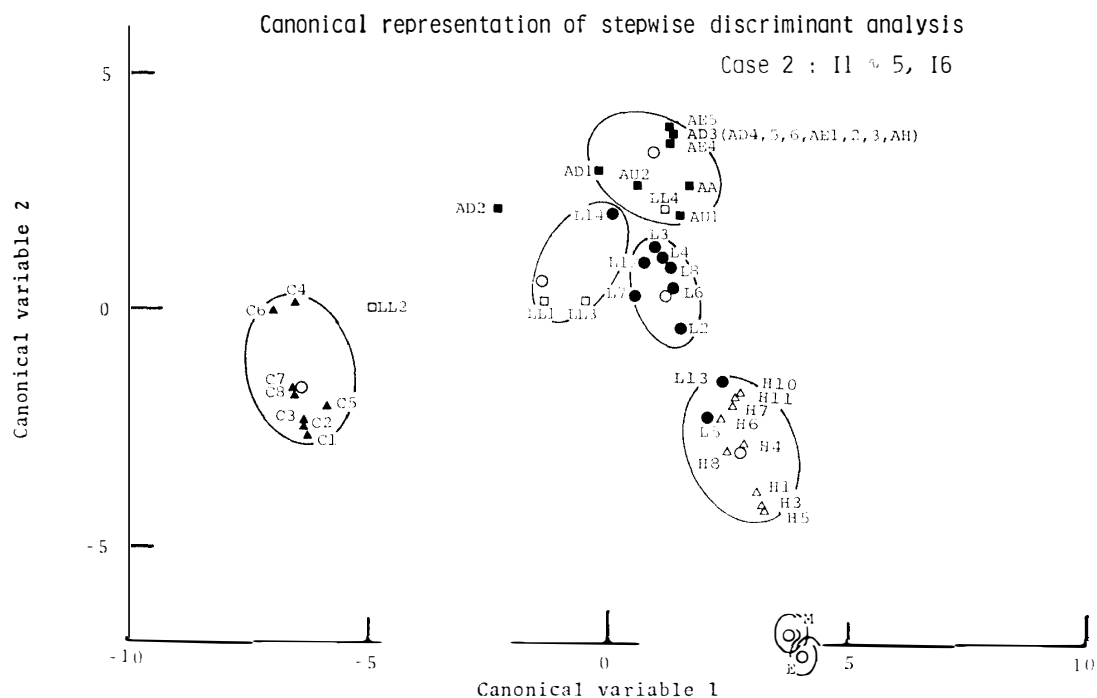


Fig. 2. Canonical representation of stepwise discriminant analysis for case 2: 11 to 15, and 16.

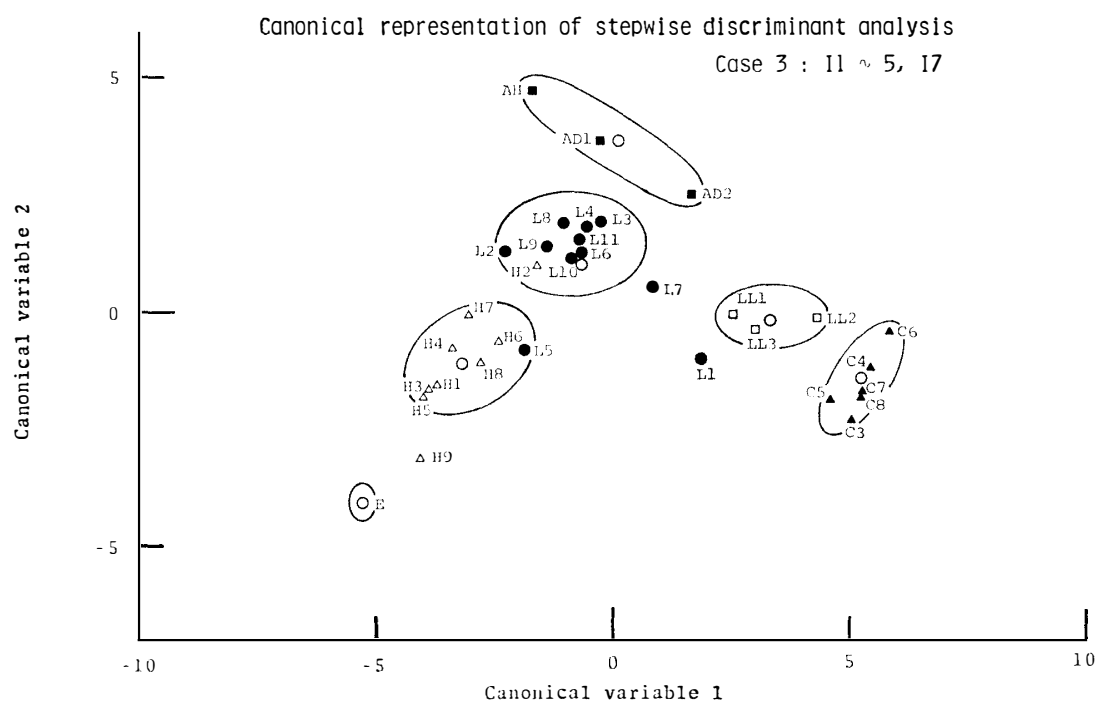


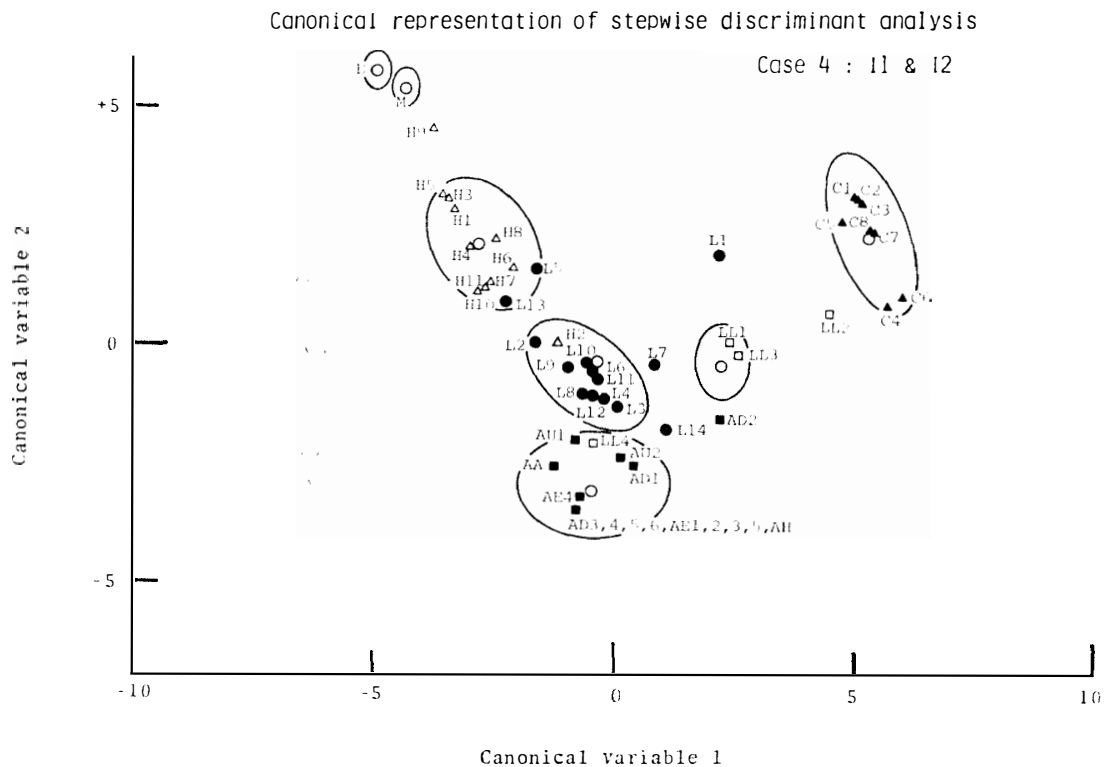
Fig. 3. Canonical representation of stepwise discriminant analysis for case 3: 11 to 15, and 17.

Hills-769 are in-between L- and LL-chondrite groups and have long amalgamation distances. L₂ Y-7304 (m) is in-between H- and L-chondrites groups by STW, but still belongs to L-chondrite groups by CLT.

LL₂ Y-74646 is in-between C- and H-chondrites groups.

4.4 Case 4: I1 and I2 (Fig. 4)

This case uses only two variables, but the results of classification by both STW and CLT are almost the same as that of case 1, except that AD₂ Y-74013 gets near LL-chondrite group.



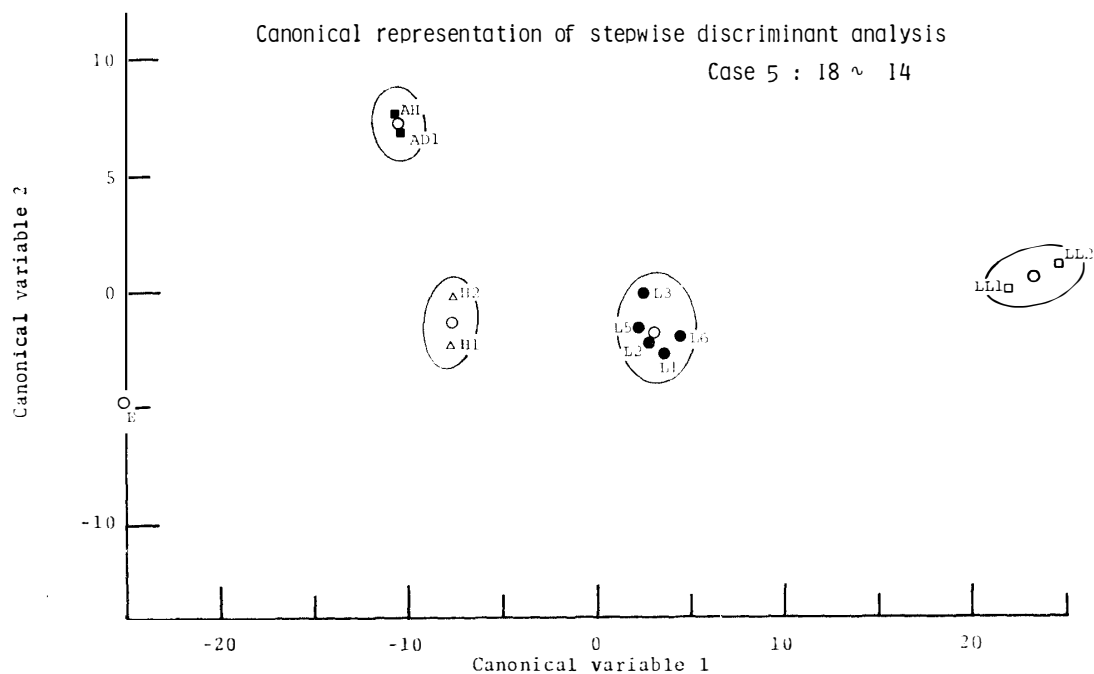


Fig. 5. Canonical representation of stepwise discriminant analysis for case 5: 18 to 114.

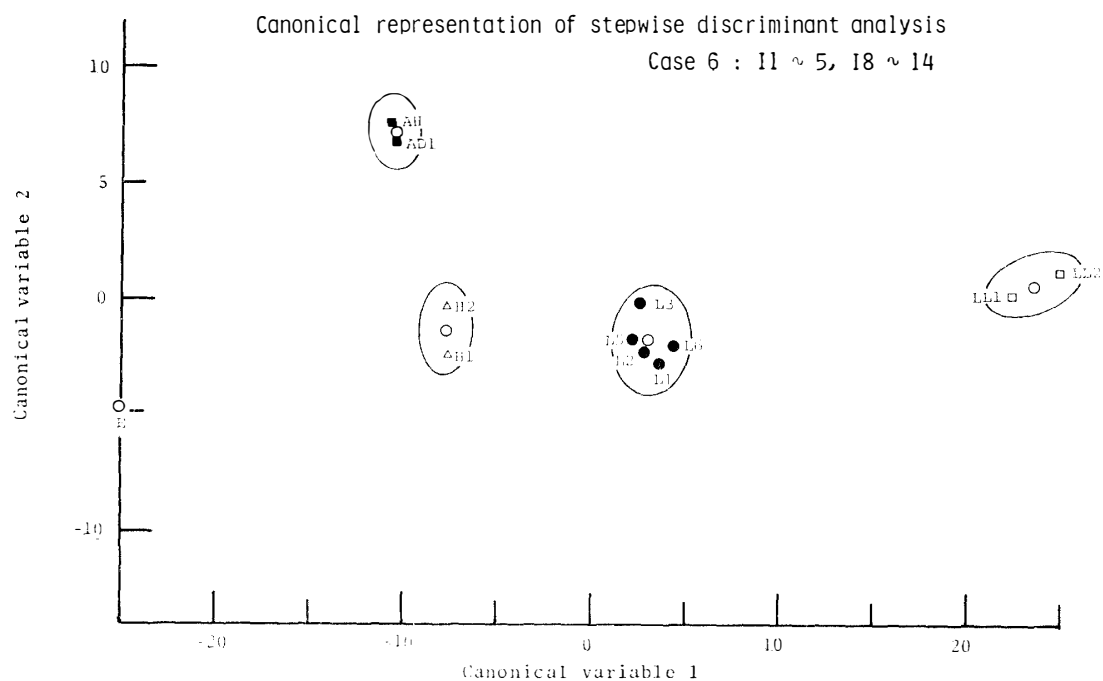
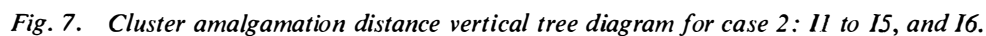


Fig. 6. Canonical representation of stepwise discriminant analysis for case 6: 11 to 15 and 18 to 114.

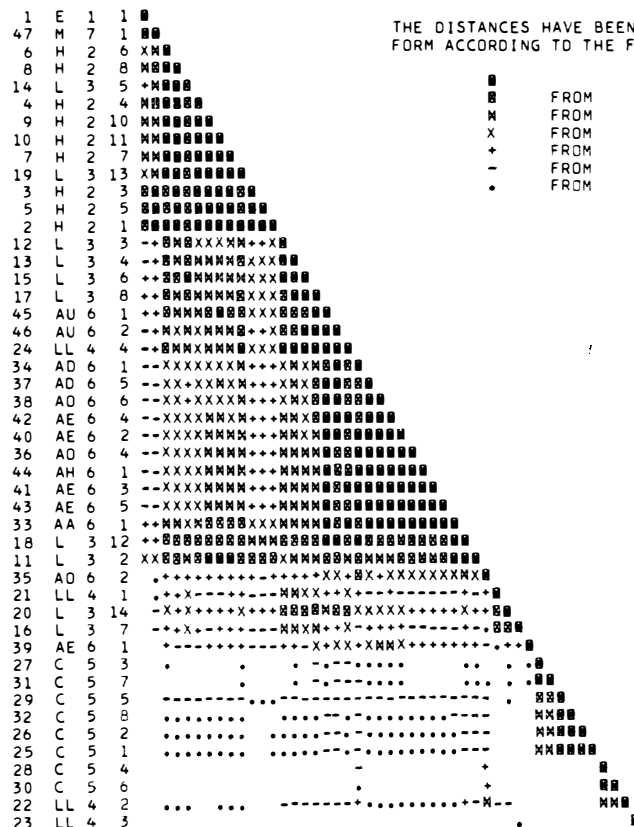


Two kinds of multivariate statistical analyses are applied to classification of the total 58 antarctic stony meteorites. Two methods, *i.e.* a cluster analysis and a step-wise discriminant analysis, give basically the same classification except for a few differences.

As the previous study by NAGATA (1979a, b) made remarks on H₂ Y-7301 (i)

DISTANCES BETWEEN CASES REPRESENTED IN SHADED FORM.
HEAVY SHADING INDICATES SMALL DISTANCES.

CASE CASE
NO. LABEL



THE DISTANCES HAVE BEEN REPRESENTED ABOVE IN SHADED
FORM ACCORDING TO THE FOLLOWING SCHEME

| | | | |
|---|------|--------------|-------|
| ■ | FROM | LESS THAN | 1.077 |
| ■ | FROM | TO | 1.629 |
| ■ | FROM | TO | 2.043 |
| ■ | FROM | TO | 2.539 |
| ■ | FROM | TO | 3.533 |
| ■ | FROM | TO | 4.113 |
| ■ | FROM | TO | 4.361 |
| ■ | FROM | GREATER THAN | 4.361 |

Fig. 8. Shaded form of amalgamation distance tree diagram for case 2: 11 to 15, and 16. Gray levels of shading are shown in the figure.

sample that the metallic phases of this chondrite are strongly weathered and metals are oxidized or hydro-oxidized, the present study also shows that H₂ belongs to the L-chondrite group near L₉ Bjurböle, although the stepwise discriminant analysis with the chemical and petrographical components indicates H₂ Y-7301 (j) is an H chondrite.

Magnetic characteristics including θ_s of H₉ Mt. Brown are in-between H and E chondrites L₅ Fukutomi with magnetic components belongs to H-chondrite group near H₅ Yonozu in cases 1 and 2, whereas the cluster analysis of case 3 shows this sample belongs to its own L-chondrite group. L₁₃ Y-74354 is also belonging to H-chondrite group near H₁₀ Y-74054, H₁₁ Y-74115, and H₇ Seminole. It is, however, not elucidated physically that L chondrite belongs to H chondrite, even taking account of weathering or oxidizing. Among samples described in the previous section, there are several which exist in-between their own groups and other groups depending on the analysis methods and the cases. One of partly possible reasons is that some stony

meteorites are generally brecciated and the pieces from a sample are different ones for magnetic and petrographical tests.

Chemical and petrographical properties as shown in case 5 classify the samples into their own groups with good separation except for H₂ Y-7301 (j) belonging to L-chondrite group by the cluster analysis.

That the combination of case 1 (I1 to I5) and case 5 (I8 to I14) shows almost the same result as case 5 implies that the effects of chemical and petrographical properties on the classification dominate over the effects of magnetic components.

A more important fact is that even though case 4 uses only two variables, I_s and $I_s(\alpha)/I_s$, the result shows almost identical classification as derived from cases 1, 2 and 3. This means I_s and $I_s(\alpha)/I_s$ are good enough parameters to classify the antarctic stony meteorites as has been pointed out by NAGATA (1979b).

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